Note on the Retarded van der Waals Potential within the Dipole Approximation

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Abstract. We examine the dipole approximated Pauli–Fierz Hamiltonians of the nonrelativistic QED. We assume that the Coulomb potential of the nuclei together with the Coulomb interaction between the electrons can be approximated by harmonic potentials. By an exact diagonalization method, we prove that the binding energy of the two hydrogen atoms behaves as R^{-7} , provided that the distance between atoms R is sufficiently large. We employ the Feynman's representation of the quantized radiation fields which enables us to diagonalize Hamiltonians, rigorously. Our result supports the famous conjecture by Casimir and Polder.

Key words: retarded van der Waals potential; non-relativistic QED; Pauli–Fierz Hamiltonian; dipole approximation

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1 Introduction

London was the first to explain attractive interactions between neutral atoms or molecules by applying quantum mechanics [16]. Nowadays, the attractive forces are called the van der Waals–London forces, and are described by the potential energy decaying as R^{-6} for R sufficiently large.¹ Here, R denotes the distance between two atoms or molecules. It is recognized that these forces come from the quantum fluctuations of the charges inside the atoms. Because even a simple hydrogen atom displays a fluctuating dipole, the van der Waals–London forces are ubiquitous and therefore very fundamental.

Casimir and Polder took the interactions between electrons and the quantized radiation fields into consideration and performed the fourth order perturbative computations [6]. They found that the finiteness of the speed of light weakens the correlation between nearby dipoles and causes the attractive potential between atoms to behave as

$$V_{\rm CP}(R) \cong -\frac{23}{4\pi} \left(\frac{1}{2\pi}\right)^2 \frac{1}{R^7} \alpha_A \alpha_B, \qquad R \gg 1, \tag{1.1}$$

where α_A and α_B are the static polarizability of the atoms. The potential $V_{\rm CP}$ is called the Casimir–Polder potential or the retarded van der Waals potential. For reviews, see, e.g., [5, 11, 13, 18, 19]. Although this result is plausible, Casimir–Polder's arguments are heuristic, and lack mathematical rigor.

¹More precisely, if one takes the interactions between electrons and the quantized Maxwell field according to non-relativistic QED into account, the R^{-6} behavior is true for the near-field region (very vaguely "sufficiently large R but not too large", and discussed on [7, p. 157] and [21]), but for the far-field region (where "retardation effects become important") the presented results show a R^{-7} behavior. In the approximation where the quantum fluctuations of the Maxwell field are ignored, only the electrostatic Coulomb interaction remains. In this case, the binding energy behaves as R^{-6} provided that R is sufficiently large. This R^{-6} behavior is well-understood, mathematically [1, 2, 15, 22].

There are few rigorous results concerning the Casimir–Polder potential; In [20, 21], Miyao and Spohn gave a path integral formula for $V_{\rm CP}$ and applied it to computing the second cumulant. Under the assumption that all of higher order cumulants behave as $O(R^{-9})$ and their coefficients are small enough to control, they rigorously refound that $V_{\rm CP}$ behaves as R^{-7} as $R \to \infty$. Although this assumption appears to be plausible, to prove it is extremely hard. Therefore, to give a mathematical foundation of the Casimir–Polder potential is an open problem even today.

In the present paper, we will examine the Pauli–Fierz model under the following assumptions [24, equations (13.127) and (13.123)]:

- (C.1) the dipole approximation (see (2.2));
- (C.2) the electrons are strongly bound around each nucleus (see (2.3) and (2.4)).

The dipole approximation (C.1) is widely accepted as a convenient procedure in the community of the nonrelativistic QED [24]. The assumption (C.2) is often useful when we study the low energy behavior of the system. Under the assumptions, we prove that the binding energy for two hydrogen atoms actually behaves as R^{-7} . In the context of the Born–Oppenheimer approximation, this indicates that the effective potential between two hydrogen atoms behaves as R^{-7} too. This result supports our assumptions for the model without dipole approximation, and is expected to become a starting point for study of the non-approximated model. Our proof relies on the fact that the dipole approximated Hamiltoninas can be diagonalized by applying Feynman's representation of the quantized radiation fields [8]. It has been believed that the dipole approximated model also exhibits R^{-7} behavior by the forth order perturbation theory. However, the arguments concering the error terms are completely missing. Indeed, this part is tacitly assumed to be trivial in literatures. In this paper, we actually perform systematic error estimates which are far from trivial.

In mathematical physics, it is known that rigorous studies of the Pauli–Fierz Hamiltonian require an extra care due to the infamous infrared problem [4, 10, 24]. Fortunetely, within the assumptions (C.1) and (C.2), we can control the problem relatively easily.

Before we proceed, we have additional remarks. In his Ph.D. Thesis [12], Koppen studied the retarded van der Waals potential; he examined the Pauli–Fierz model with the dipole approximation (C.1), but the condition (C.2) is not assumed in [12]. In contrast to the present study, he imposed the infrared cutoff σ on the Hamiltonian in order to apply the naive perturbation theory and obtained an expansion formula for the binding energy: $E_{\sigma}(R) = \sum_{i=0}^{\infty} e^i V_i^{\sigma}(R)$. Then he removed the infrared cutoff from each term: $V_i(R) := \lim_{\sigma \to +0} V_i^{\sigma}(R)$. Finally, he proved that some $V_i(R)$ satisfies (1.1). His observation could be regareded as a nice starting point of mathematical analysis of the retarded van der Waals potential, however, there are still some problems to be considered. For example, the magnetic contributions to the $-R^{-7}$ decay are completely overlooked. In addition, in the mathematical study of the Pauli–Fierz model, it is well-known that to prove that $\lim_{\sigma \to +0} E_{\sigma}(R) = \sum_{i=0}^{\infty} e^i \lim_{\sigma \to +0} V_i^{\sigma}(R)$ is very hard problem, the aforementioned infrared problem.

Our contributions are

- to provide a minimal QED model which can rigorously explain the Casimir–Polder potential by a relatively simple and easy way;
- to perform systematic error estimates without the infrared cutoff.

In this way, the present paper and the thesis [12] are complementary to each other.

Since the electrons obey Fermi–Dirac statistics, the wave functions of the two-electron system belong to $(\mathfrak{H} \wedge \mathfrak{H}) \otimes \mathfrak{F}(L^2(\mathbb{R}^3 \times \{1, 2\}))$, where $\mathfrak{H} = L^2(\mathbb{R}^3) \otimes \mathbb{C}^2$, the Hilbert space with spin 1/2,

the symbol \wedge indicates the anti-symmetric tensor product and $\mathfrak{F}(L^2(\mathbb{R}^3 \times \{1,2\}))$ is the Fock space over $L^2(\mathbb{R}^3 \times \{1,2\})$. Usually, the ground state of this system is a spin singlet. Thus, the spatial part of the ground state is symmetric and we can end up with minimizing the energy in an unrestricted manner on $(L^2(\mathbb{R}^3) \otimes L^2(\mathbb{R}^3)) \otimes \mathfrak{F}(L^2(\mathbb{R}^3 \times \{1,2\}))$. For this reason, we perform our analysis on $(L^2(\mathbb{R}^3) \otimes L^2(\mathbb{R}^3)) \otimes \mathfrak{F}(L^2(\mathbb{R}^3 \times \{1,2\}))$.² However, it should be mentioned that our observation here can not be extended to general N-electron systems, directly.

In fairness, we mention the following two difficulties of the assumptions (C.1) and (C.2). For details, see discussions in Section 9.

- The condition (C.2) breaks the indistinguishability of the electrons.
- Under the conditions (C.1) and (C.2), we cannot reproduce the exact cancellation of the term with R^{-6} decay (the van der Waals–London potential) by the contribution from the quantized Maxwell field. Note that this cancellation is known to be fundamental to explain the retarded van der Waals potential [20, 21].

The present paper is organized as follows. In Section 2, we introduce the dipole approximated Pauli–Fierz Hamiltonian and state the main result. In Section 3, we switch to the Feynman representation of the quantized radiation fields. This representation enables us to diagonalize the Hamiltonians as we will see in the following sections. Further, we introduce a canonical transformation which induces the quantized displacement fields in the Hamiltonians in Section 4. Section 5 is devoted to the finite volume approximation, which is a standard method in the study of the quantum field theory [3, 9]. Then we diagonalize the Hamiltonians in Sections 6 and 7. In Section 8, we give a proof of the main theorem. Section 9 is devoted to the discussions of the approximations (C.1) and (C.2). In Appendices A, B and C, we collect various auxiliary results which are needed in the main sections.

2 Main result

Let us consider a single hydrogen atom with an infinitely heavy nucleus located at the origin 0. The nonrelativistic QED Hamiltonian for this system is given by

$$H_{1e} = \frac{1}{2} \left(-i\nabla - eA(x) \right)^2 - e^2 V(x) + H_{f}.$$

The nucleus has charge e > 0, and the electron has charge -e. We assume that the charge distribution ρ satisfies the following properties:

(A.1) ϱ is normalized: $\int_{\mathbb{R}^3} \mathrm{d}x \, \varrho(x) = 1$.

(A.2) $\rho(x) = \rho(-x)$. Thus the Fourier transformation $\hat{\rho}$ is real.

(A.3) $\hat{\varrho}$ is rotation invariant, $\hat{\varrho}(k) = \hat{\varrho}_{rad}(|k|)$, of rapid decrease and smooth.

The smeared Coulomb potential V is given by

$$V(x) = \int_{\mathbb{R}^3} \mathrm{d}k \,\hat{\varrho}(k)^2 |k|^{-2} \mathrm{e}^{-\mathrm{i}k \cdot x}.$$

The photon annihilation operator is denoted by $a(k, \lambda)$. As usual, this operator satisfies the standard commutation relation:

$$[a(k,\lambda),a(k',\lambda')^*] = \delta_{\lambda\lambda'}\delta(k-k').$$

 $^{^{2}}$ Or we could simply say that one considers the "distinguishable particles", see Section 9 for detail.



Figure 1.

The quantized vector potential A(x) is defined by

$$A(x) = \sum_{\lambda=1,2} \int_{\mathbb{R}^3} \mathrm{d}k \, \frac{\hat{\varrho}(k)}{\sqrt{2|k|}} \varepsilon(k,\lambda) \big(\mathrm{e}^{-\mathrm{i}k \cdot x} a(k,\lambda)^* + \mathrm{e}^{\mathrm{i}k \cdot x} a(k,\lambda) \big),$$

where $\varepsilon(k,\lambda) = (\varepsilon_1(k,\lambda), \varepsilon_2(k,\lambda), \varepsilon_3(k,\lambda)), \lambda = 1, 2$ are polarization vectors. For concreteness, we choose as

$$\varepsilon(k,1) = \frac{(k_2, -k_1, 0)}{\sqrt{k_1^2 + k_2^2}}, \qquad \varepsilon(k,2) = \frac{k}{|k|} \wedge \varepsilon(k,1).$$
(2.1)

Note that A(x) is essentially self-adjoint. We will denote its closure by the same symbol. The field energy $H_{\rm f}$ is given by

$$H_{\rm f} = \sum_{\lambda=1,2} \int_{\mathbb{R}^3} \mathrm{d}k \, |k| a(k,\lambda)^* a(k,\lambda).$$

The operator H_{1e} acts in the Hilbert space $L^2(\mathbb{R}^3) \otimes \mathfrak{F}(L^2(\mathbb{R}^3 \times \{1,2\}))$, where $\mathfrak{F}(\mathfrak{h})$ is the bosonic Fock space over \mathfrak{h} : $\mathfrak{F}(\mathfrak{h}) = \bigoplus_{n=0}^{\infty} \mathfrak{h}^{\otimes_{\mathbb{S}} n}$. Here, $\otimes_{\mathbb{S}}$ indicates the symmetric tensor product.

To examine the Casimir–Polder potential, we consider two hydrogen atoms, one located at the origin and the other at r = (0, 0, R) with R > 0. For computational convenience, we define the position of the second electron relative to r, see Fig. 1. Then the two-electron Hamiltonian reads

$$H_{2e} = \frac{1}{2} \left(-i\nabla_1 - eA(x_1) \right)^2 - e^2 V(x_1) + \frac{1}{2} \left(-i\nabla_2 - eA(x_2 + r) \right)^2 - e^2 V(x_2) + e^2 V_R(x_1, x_2) + H_f$$

with

$$V_R(x_1, x_2) = -V(x_1 - r) - V(x_2 + r) + V(r) + V(r + x_2 - x_1)$$

= $\int_{\mathbb{R}^3} dk \,\hat{\varrho}(k)^2 |k|^{-2} (1 - e^{-ik \cdot x_1}) (1 - e^{ik \cdot x_2}) e^{ik \cdot r}.$

The operator H_{2e} acts in $L^2(\mathbb{R}^3_{x_1}) \otimes L^2(\mathbb{R}^3_{x_2}) \otimes \mathfrak{F}(L^2(\mathbb{R}^3_k \times \{1,2\})).$

The dipole approximation (C.1) means the following replacement:

$$A(x_1) \rightsquigarrow A(0), \qquad A(x_2+r) \rightsquigarrow A(r).$$
 (2.2)

By the assumption (C.2), we can take x_1 and x_2 sufficiently small. Therefore, we assume that the Coulomb potential of the nuclei together with the Coulomb interaction between the electrons can be approximated by harmonic potentials. Then one has

$$V(x_j) \simeq -\frac{1}{2}\nu_0^2 x_j^2 + \text{const}$$
 (2.3)

with $\nu_0^2 = \frac{1}{3} \int \mathrm{d}k \,\hat{\varrho}(k)^2$ and

$$V_R(x_1, x_2) \simeq \int \mathrm{d}k \,\hat{\varrho}(k)^2 \mathrm{e}^{\mathrm{i}k \cdot r} \left(x_1 \cdot \hat{k}\right) \left(x_2 \cdot \hat{k}\right) \tag{2.4}$$

with $\hat{k} = k/|k|$. Hence, we arrive at

$$H_{\rm D1e} = \frac{1}{2} \left(-i\nabla - eA(0) \right)^2 + \frac{1}{2} e^2 \nu_0^2 x^2 + H_{\rm f}$$

and

$$\begin{aligned} H_{\rm D2e} &= \frac{1}{2} \Big(-\mathrm{i} \nabla_1 - eA(0) \Big)^2 + \frac{1}{2} e^2 \nu_0^2 x_1^2 + \frac{1}{2} \Big(-\mathrm{i} \nabla_2 - eA(r) \Big)^2 + \frac{1}{2} e^2 \nu_0^2 x_2^2 \\ &+ e^2 \int_{\mathbb{R}^3} \mathrm{d} k \, \hat{\varrho}(k)^2 \mathrm{e}^{\mathrm{i} k \cdot r} (x_1 \cdot \hat{k}) (x_2 \cdot \hat{k}) + H_{\rm f}. \end{aligned}$$

Note that H_{D1e} and H_{D2e} are self-adjoint and bounded from below [14], because the cross-term $\int dk \,\hat{\varrho}(k)^2 e^{ik \cdot r} (x_1 \cdot \hat{k}) (x_2 \cdot \hat{k})$ becomes very small provided that R is large enough. As for physical discussions of the approximation above, see Section 9 in detail.

In what follows, we assume an additional condition:

(A.4) We regard ν_0 as a parameter. Thus, ν_0 is independent of ρ .

Hence, there are three parameters e, R and ν_0 in our models.

Theorem 2.1. Let $E(R) = \inf \operatorname{spec}(H_{D2e})$ and let $E = \inf \operatorname{spec}(H_{D1e})$, where $\operatorname{spec}(X)$ indicates the spectrum of a linear operator X. Let

$$c_{\infty} = \max\left\{\sqrt{2}e\left\||k|^{-1}\hat{\varrho}\right\|_{L^{2}}, \frac{\||k|\hat{\varrho}\|_{L^{2}}}{\sqrt{2}e\nu_{0}^{2}}, \frac{\|\hat{\varrho}\|_{L^{2}}}{\nu_{0}}\right\}.$$

Choose e and ν_0 such that $c_{\infty} < 1/2, 1 \leq \sqrt{2}e\nu_0$ and $\sqrt{2}e\|\hat{\varrho}\|_{L^2} < 1$. Then one has

$$\lim_{R \to \infty} R^7 (2E - E(R)) = \frac{23}{4\pi} \left(\frac{1}{2\pi}\right)^2 \left(\frac{1}{4}\alpha_{\rm E,at}\right)^2,$$

where $\alpha_{\rm E,at} = \nu_0^{-2}$.

Remark 2.2.

1. The constant $\alpha_{E,at}$ is the dipole moment of a decoupled atom, i.e.,

$$\alpha_{\rm E,at} = \frac{2}{3} \langle \psi_{\rm at} | x \cdot (h_{\rm at} - 3e\nu_0/2)^{-1} x \psi_{\rm at} \rangle, \qquad (2.5)$$

where $h_{\rm at} = -\frac{1}{2}\Delta + \frac{e^2\nu_0^2}{2}x^2$ and $\psi_{\rm at}$ is the ground state of $h_{\rm at}$. Note that $x_j\psi_{\rm at}$ is orthogonal to $\psi_{\rm at}$: $\langle \psi_{\rm at} | x_j\psi_{\rm at} \rangle = 0$. Thus, the vectors $(h_{\rm at} - 3e\nu_0/2)^{-1}x_j\psi_{\rm at}$ in (2.5) are mathematically meaningful.

2. The restrictions of the parameters in Theorem 2.1 come from technical reasons: As we will see in the later sections, these are needed in order to control the perturbative expansions for E and E(R).

Example 2.3. Let $\eta \in \mathscr{S}(\mathbb{R}^3)$, the Schwartz space. Suppose that η satisfies the following:

- $\eta(0) = (2\pi)^{-3/2};$
- $\eta(k)$ is real-valued;
- $\eta(k) = \eta_{\mathrm{rad}}(|k|).$

For given $\xi > 0$, we define ρ by

$$\hat{\varrho}(k) = \eta(\xi k)$$

Then ρ satisfies (A.1)–(A.3). In addition, since

 $\|\hat{\varrho}\|_{L^2} \propto \xi^{-3/2}, \qquad \||k|\hat{\varrho}\|_{L^2} \propto \xi^{-5/2}, \qquad \||k|^{-1}\hat{\varrho}\|_{L^2} \propto \xi^{-1/2},$

the all assumptions in Theorem 2.1 are fulfilled, provided that ξ is large enough. Note that a typical choice of η is $\eta(k) = (2\pi)^{-3/2} e^{-|k|^2}$.

3 Feynman Hamiltonians

3.1 Preliminaries

To prove our main result, let us introduce Feynman Hamiltonians of the nonrelativistic QED [8]. These Hamiltonians can be diagonalized readily as we will see in Sections 6 and 7.

First, remark the following identification:

$$L^2(\mathbb{R}^3) = L^2_e(\mathbb{R}^3) \oplus L^2_o(\mathbb{R}^3),$$

where

$$L_e^2(\mathbb{R}^3) = \{ f \in L^2(\mathbb{R}^3) | f(-x) = f(x) \text{ a.e.} \}, L_o^2(\mathbb{R}^3) = \{ f \in L^2(\mathbb{R}^3) | f(-x) = -f(x) \text{ a.e.} \}.$$

For notational convenience, we denote by $\varepsilon_j(\cdot, \lambda)$ the multiplication operator by the function $\varepsilon_j(\cdot, \lambda)$.

We begin with the following lemma.

Lemma 3.1. Let

$$\begin{split} \mathfrak{H}_1 &= \bigcup_{j=1}^3 \overline{\operatorname{ran}}\big(\varepsilon_j(\cdot,1) \upharpoonright L_e^2\big(\mathbb{R}^3\big)\big), \qquad \mathfrak{H}_2 = \bigcup_{j=1}^3 \overline{\operatorname{ran}}\big(\varepsilon_j(\cdot,2) \upharpoonright L_e^2\big(\mathbb{R}^3\big)\big), \\ \mathfrak{H}_3 &= \bigcup_{j=1}^3 \overline{\operatorname{ran}}\big(\varepsilon_j(\cdot,1) \upharpoonright L_o^2\big(\mathbb{R}^3\big)\big), \qquad \mathfrak{H}_4 = \bigcup_{j=1}^3 \overline{\operatorname{ran}}\big(\varepsilon_j(\cdot,2) \upharpoonright L_o^2\big(\mathbb{R}^3\big)\big), \end{split}$$

where $A \upharpoonright \mathfrak{X}$ indicates the restriction of A to \mathfrak{X} . Then $\mathfrak{H}_1, \mathfrak{H}_2, \mathfrak{H}_3$ and \mathfrak{H}_4 are subspaces of $L^2(\mathbb{R}^3)$.

Proof. Let $\mathbb{D} = \{k \in \mathbb{R}^3 | k_1 \neq 0, k_2 \neq 0, k_3 \neq 0\}$. Trivially, $\varepsilon_j(k, \lambda)$ is well-defined on \mathbb{D} . In addition, $\varepsilon_1(k, 1)^{-1}$, $\varepsilon_2(k, 1)^{-1}$, $\varepsilon_1(k, 2)^{-1}$, $\varepsilon_2(k, 2)^{-1}$ and $\varepsilon_3(k, 2)^{-1}$ are well-defined on \mathbb{D} .³ Let $C_0(\mathbb{D})$ be the set of continuous functions on \mathbb{D} of compact support. Because the Lebesgue measure of \mathbb{D}^c , the complement of \mathbb{D} , is equal to zero, $C_0(\mathbb{D})$ is dense in $L^2(\mathbb{R}^3)$. Thus, it holds that

$$\overline{\operatorname{ran}}(\varepsilon_j(\cdot,1) \upharpoonright L^2_e(\mathbb{R}^3)) = \overline{\operatorname{ran}}(\varepsilon_j(\cdot,1) \upharpoonright C_{0,e}(\mathbb{D})), \qquad j = 1, 2,$$
(3.1)

where $C_{0,e}(\mathbb{D}) = \{ f \in C_0(\mathbb{D}) \mid f(-k) = f(k) \}.$

Let $F, G \in \mathfrak{H}_1$. Then there exist $i, j \in \{1, 2\}$ such that $F \in \overline{\operatorname{ran}}(\varepsilon_j(\cdot, 1) \upharpoonright L^2_e(\mathbb{R}^3))$ and $G \in \overline{\operatorname{ran}}(\varepsilon_i(\cdot, 1) \upharpoonright L^2_e(\mathbb{R}^3))$. By (3.1), there exist approximating sequences $(F_n) \subset \operatorname{ran}(\varepsilon_j(\cdot, 1) \upharpoonright C_{0,e}(\mathbb{D}))$ and $(G_n) \subset \operatorname{ran}(\varepsilon_i(\cdot, 1) \upharpoonright C_{0,e}(\mathbb{D}))$ such that $||F - F_n|| \to 0$ and $||G - G_n|| \to 0$ as $n \to 0$. Hence, for each $\alpha, \beta \in \mathbb{C}$, it holds that

$$\alpha F_n + \beta G_n \to \alpha F + \beta G \qquad \text{as } n \to \infty.$$
 (3.2)

Note that we can write $F_n = \varepsilon_j(\cdot, 1)f_n$ and $G_n = \varepsilon_i(\cdot, 1)g_n$ with $f_n, g_n \in C_{0,e}(\mathbb{D})$. Thus, we have $G_n = \varepsilon_j(\cdot, 1)g'_n$, where $g'_n = \varepsilon_j(\cdot, 1)^{-1}\varepsilon_i(\cdot, 1)g_n$. Because $\varepsilon_j(k, 1)^{-1}\varepsilon_i(k, 1)$ is an even function on \mathbb{D} , we see that $g'_n \in C_{0,e}(\mathbb{D})$. Accordingly, $\alpha F_n + \beta G_n = \varepsilon_j(\cdot, 1)(\alpha f_n + \beta g'_n) \in \operatorname{ran}(\varepsilon_j(\cdot, 1) \upharpoonright C_{0,e}(\mathbb{D}))$. Combining this, (3.1) and (3.2), we conclude that $\alpha F + \beta G \in \mathfrak{H}_1$, in particular, \mathfrak{H}_1 is a subspace of $L^2(\mathbb{R}^3)$.

Lemma 3.2. We have the following identifications:

$$L^{2}(\mathbb{R}^{3} \times \{1,2\}) = L^{2}(\mathbb{R}^{3}) \oplus L^{2}(\mathbb{R}^{3}) = \mathfrak{H}_{1} \oplus \mathfrak{H}_{2} \oplus \mathfrak{H}_{3} \oplus \mathfrak{H}_{4}.$$

$$(3.3)$$

Proof. The first identification in (3.3) is trivial. In what follows, we will concentrate on the proof of the second identification.

Note that the multiplication operator $\varepsilon_j(\cdot, 1)^{-1}$ (j = 1, 2) is self-adjoint and dom $(\varepsilon_j(\cdot, 1)^{-1})$ is dense in $L^2(\mathbb{R}^3)$. Because $\operatorname{ran}(\varepsilon_j(\cdot, 1)) \supseteq \operatorname{dom}(\varepsilon_j(\cdot, 1)^{-1}) \supseteq C_0(\mathbb{D})$ for j = 1, 2, we obtain $\overline{\operatorname{ran}}(\varepsilon_j(\cdot, 1)) = L^2(\mathbb{R}^3)$. For each $f \in L^2(\mathbb{R}^3)$, we set $f_e(k) = \frac{1}{2}(f(k) + f(-k))$ and $f_o(k) = \frac{1}{2}(f(k) - f(-k))$. Because $\varepsilon_j(k, 1)^2$ is an even function, we have $\langle \varepsilon_j(\cdot, 1)f_e|\varepsilon_j(\cdot, 1)f_o\rangle = 0$ for all $f \in L^2(\mathbb{R}^3)$ and j = 1, 2, which implies that $\overline{\operatorname{ran}}(\varepsilon_j(\cdot, 1) \upharpoonright L^2_e(\mathbb{R}^3)) \perp \overline{\operatorname{ran}}(\varepsilon_j(\cdot, 1) \upharpoonright L^2_o(\mathbb{R}^3))$. Since

$$\underbrace{\varepsilon_j(\cdot,1)f}_{\operatorname{Fran}(\varepsilon_j(\cdot,1))} = \underbrace{\varepsilon_j(\cdot,1)f_e}_{\operatorname{Fran}(\varepsilon_j(\cdot,1)\upharpoonright L^2_e(\mathbb{R}^3))} + \underbrace{\varepsilon_j(\cdot,1)f_o}_{\operatorname{Fran}(\varepsilon_j(\cdot,1)\upharpoonright L^2_o(\mathbb{R}^3))}, \qquad f \in L^2(\mathbb{R}^3),$$

we conclude that

e

$$L^{2}(\mathbb{R}^{3}) = \overline{\operatorname{ran}}(\varepsilon_{j}(\cdot, 1) \upharpoonright L^{2}_{e}(\mathbb{R}^{3})) \oplus \overline{\operatorname{ran}}(\varepsilon_{j}(\cdot, 1) \upharpoonright L^{2}_{o}(\mathbb{R}^{3}))$$

for j = 1, 2.

For $i, i' \in \{1, 2\}$, we set $\mu_{ii'}^{(1)}(k) = \varepsilon_i(k, 1)\varepsilon_{i'}(k, 1)$. Because $\mu_{ii'}^{(1)}(k)$ is an even function, we see that, for each $f \in L^2_e(\mathbb{R}^3)$ and $g \in L^2_o(\mathbb{R}^3)$,

$$\langle \varepsilon_i(\cdot, 1)f|\varepsilon_{i'}(\cdot, 1)g\rangle = \langle f|\mu_{ii'}^{(1)}g\rangle = 0.$$

Therefore, $\mathfrak{H}_1 \perp \mathfrak{H}_3$ holds. Because $\mathfrak{H}_1 \oplus \mathfrak{H}_3 \supseteq \overline{\operatorname{ran}}(\varepsilon_j(\cdot, 1) \upharpoonright L^2_e(\mathbb{R}^3)) \oplus \overline{\operatorname{ran}}(\varepsilon_j(\cdot, 1) \upharpoonright L^2_o(\mathbb{R}^3))$, we finally arrive at $L^2(\mathbb{R}^3) = \mathfrak{H}_1 \oplus \mathfrak{H}_3$. By arguments similar to the above, we get that $L^2(\mathbb{R}^3) = \mathfrak{H}_2 \oplus \mathfrak{H}_4$.

We will construct a useful identification between $\mathfrak{F}(L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3))$ and $\bigotimes_{\lambda=1}^4 \mathfrak{F}(\mathfrak{H}_{\lambda})$ in Section 3.4. For this purpose, we recall some basic definitions in Sections 3.2 and 3.3.

³These facts immediately follow from (2.1). Here, note that $\varepsilon(k, 2)$ is written as $\varepsilon(k, 2) = (k_1k_3, k_2k_3, -k_1^2 - k_2^2)/|k|\sqrt{k_1^2 + k_2^2}$.

3.2 Second quantized operators in $\mathfrak{F}(L^2(\mathbb{R}^3)\oplus L^2(\mathbb{R}^3))$

Let $a(f_1 \oplus f_2)$ be the annihilation operator acting in $\mathfrak{F}(L^2(\mathbb{R}^3 \times \{1,2\})) = \mathfrak{F}(L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3))$. As usual, we express this operator as

$$a(f_1 \oplus f_2) = \sum_{\lambda=1,2} \int_{\mathbb{R}^3} \mathrm{d}k \, f_{\lambda}^*(k) a(k,\lambda).$$

The Fock vacuum in $\mathfrak{F}(L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3))$ is denoted by Ψ_0 . Let F be a real-valued function on \mathbb{R}^3 which is finite almost everywhere. The multiplication operator by F is also written as F. The second quantization of $F \oplus F$ is then given by

$$\mathrm{d}\Gamma(F\oplus F) = 0 \oplus \left[\bigoplus_{n=1}^{\infty}\sum_{j=1}^{n}1\otimes\cdots\otimes\underbrace{(F\oplus F)}_{j^{\mathrm{th}}}\otimes\cdots\otimes1\right].$$

Needless to say, $d\Gamma(F \oplus F)$ acts in $\mathfrak{F}(L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3))$. It is known that $d\Gamma(F \oplus F)$ is essentially self-adjoint on a dense subspace

$$\left\{\Psi = \{\Psi_n\}_{n=0}^{\infty} \,|\, \Psi_n \in (\operatorname{dom}(F) \oplus \operatorname{dom}(F))^{\odot n}, \exists N \in \mathbb{N} \text{ s.t. } \Psi_m = 0 \;\forall \, m > N\right\}$$

where \odot indicates the algebraic tensor product. We will denote the closure of $d\Gamma(F \oplus F)$ by the same symbol. Symbolically, we express $d\Gamma(F \oplus F)$ as

$$\mathrm{d}\Gamma(F\oplus F) = \sum_{\lambda=1,2} \int_{\mathbb{R}^3} \mathrm{d}k \, F(k) a(k,\lambda)^* a(k,\lambda).$$

3.3 Second quantized operators in $\bigotimes_{\lambda=1}^{4} \mathfrak{F}(\mathfrak{H}_{\lambda})$

Let $a_{\lambda}(f_{\lambda})$ be the annihilation operator on $\mathfrak{F}(\mathfrak{H}_{\lambda})$. We employ the following identifications: $a_1(f_1) = a_1(f_1) \otimes \mathbb{1} \otimes \mathbb{1} \otimes \mathbb{1}, a_2(f_2) = \mathbb{1} \otimes a_2(f_2) \otimes \mathbb{1} \otimes \mathbb{1} \otimes \mathbb{1}$ and so on. Thus, $a_{\lambda}(f_{\lambda})$ can be regarded as a linear operator acting in the Hilbert space $\bigotimes_{\lambda=1}^{4} \mathfrak{F}(\mathfrak{H}_{\lambda})$. Let F be a real-valued function on \mathbb{R}^3 . Suppose that F is even: F(-k) = F(k) a.e., $d\Gamma_{\lambda}(F)$ denotes the second quantization of F which acts in $\mathfrak{F}(\mathfrak{H}_{\lambda})$. As before, we can also regard $d\Gamma_{\lambda}(F)$ as a linear operator acting in $\bigotimes_{\lambda=1}^{4} \mathfrak{F}(\mathfrak{H}_{\lambda})$. The Fock vacuum in $\mathfrak{F}(\mathfrak{H}_{\lambda})$ is denoted by Ψ_{λ} . We will freely use the following notations:

$$a_{\lambda}(f) = \int_{\mathbb{R}^3} \mathrm{d}k \, f(k)^* a(k,\lambda), \qquad f \in \mathfrak{H}_{\lambda},$$
$$\mathrm{d}\Gamma_{\lambda}(F) = \int_{\mathbb{R}^3} \mathrm{d}k \, F(k) a^*(k,\lambda) a(k,\lambda), \qquad \lambda = 1, 2, 3, 4$$

3.4 Identifications between $\mathfrak{F}(L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3))$ and $\bigotimes_{\lambda=1}^4 \mathfrak{F}(\mathfrak{H}_\lambda)$

For each $\boldsymbol{f} = (f_1, f_2) \in L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3)$, we set

$$b_{ij}(\boldsymbol{f}) = a(\varepsilon_i(\cdot, 1)f_1 \oplus \varepsilon_j(\cdot, 2)f_2),$$

$$c_{ij}(\boldsymbol{f}) = a_1(\varepsilon_i(\cdot, 1)f_{1,e}) + a_2(\varepsilon_j(\cdot, 2)f_{2,e}) - ia_3(\varepsilon_i(\cdot, 1)f_{1,o}) - ia_4(\varepsilon_j(\cdot, 2)f_{2,o}),$$
(3.4)

where $f_e(k) = (f(k) + f(-k))/2$ and $f_o(k) = (f(k) - f(-k))/2$. Let Ψ_0 be the Fock vacuum in $\mathfrak{F}(L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3)) : \Psi_0 = 1 \oplus 0 \oplus 0 \oplus \cdots$.

Lemma 3.3. We define a linear operator $V: \mathfrak{F}(L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3)) \to \bigotimes_{\lambda=1}^4 \mathfrak{F}(\mathfrak{H}_{\lambda})$ by

$$V\Psi_{0} = \bigotimes_{\lambda=1}^{4} \Psi_{\lambda},$$
$$V\left[\prod_{\ell=1}^{N} b_{i_{\ell}j_{\ell}}(\boldsymbol{f}_{\ell})^{*}\right]\Psi_{0} = \left[\prod_{\ell=1}^{N} c_{i_{\ell}j_{\ell}}(\boldsymbol{f}_{\ell})^{*}\right]\bigotimes_{\lambda=1}^{4} \Psi_{\lambda}$$

for each $f_1, \ldots, f_N \in L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3)$ and $N \in \mathbb{N}$. Then V can be extended to the unitary operator. In what follows, we denote the extension by the same symbol. Then we have

$$Vb_{ij}(\boldsymbol{f})V^{-1} = \overline{c_{ij}(\boldsymbol{f})}$$
(3.5)

for each $\mathbf{f} \in L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3)$ and $i, j \in \{1, 2, 3\}$, where the bar indicates the closure of the operator.

Proof. For $i, i' \in \{1, 2, 3\}$, we set

$$\mu_{ii'}^{(1)}(k) = \varepsilon_i(k,1)\varepsilon_{i'}(k,1), \qquad \mu_{ii'}^{(2)}(k) = \varepsilon_i(k,2)\varepsilon_{i'}(k,2).$$

For $\boldsymbol{f}, \boldsymbol{f}' \in L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3)$ and $i, j, i', j' \in \{1, 2, 3\}$, define

$$D(\boldsymbol{f}|\boldsymbol{f}')_{ij;i'j'} = \langle f_1 | \mu_{ii'}^{(1)} f_1' \rangle + \langle f_2 | \mu_{ii'}^{(2)} f_2' \rangle.$$

First, we prove that $\{b_{ij}(\boldsymbol{f})|\boldsymbol{f} \in L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3), i, j \in \{1, 2, 3\}\}$ and $\{c_{ij}(\boldsymbol{f})|\boldsymbol{f} \in L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3), i, j \in \{1, 2, 3\}\}$ satisfy the similar commutations relations, that is,

$$[b_{ij}(f), b_{i'j'}(f')^*] = D(f|f')_{ij;i'j'}, \qquad [b_{ij}(f), b_{i'j'}(f')] = 0$$

and

$$[c_{ij}(f), c_{i'j'}(f')^*] = D(f|f')_{ij;i'j'}, \qquad [c_{ij}(f), c_{i'j'}(f')] = 0.$$

To see this, note that $\mu_{ii'}^{(1)}(k)$ and $\mu_{ii'}^{(2)}(k)$ are even functions. Thus,

$$\langle f_e | \mu_{ii'}^{(1)} g_o \rangle = 0 = \langle f_e | \mu_{ii'}^{(2)} g_o \rangle, \qquad f, g \in L^2(\mathbb{R}^3).$$

Accordingly, we have

$$[c_{ij}(\boldsymbol{f}), c_{i'j'}(\boldsymbol{f}')^*] = \langle f_{1,e} | \mu_{ii'}^{(1)} f_{1,e}' \rangle + \langle f_{2,e} | \mu_{ii'}^{(2)} f_{2,e}' \rangle + \langle f_{1,o} | \mu_{ii'}^{(1)} f_{1,o}' \rangle + \langle f_{2,0} | \mu_{ii'}^{(2)} f_{2,o}' \rangle$$

= $D(\boldsymbol{f} | \boldsymbol{f}')_{ij;i'j'}.$

To check other commutation relations are easy.

Using the above fact, we readily confirm that

$$\begin{split} \left\langle \left[\prod_{\ell=1}^{N} b_{i_{\ell} j_{\ell}}(\boldsymbol{f}_{\ell})^{*}\right] \Psi_{0} \middle| \left[\prod_{\ell=1}^{N'} b_{i_{\ell}' j_{\ell}'}(\boldsymbol{f}_{\ell}')^{*}\right] \Psi_{0} \right\rangle \\ &= \left\langle \left[\prod_{\ell=1}^{N} c_{i_{\ell} j_{\ell}}(\boldsymbol{f}_{\ell})^{*}\right] \bigotimes_{\lambda=1}^{4} \Psi_{\lambda} \middle| \left[\prod_{\ell=1}^{N'} c_{i_{\ell}' j_{\ell}'}(\boldsymbol{f}_{\ell}')^{*}\right] \bigotimes_{\lambda=1}^{4} \Psi_{\lambda} \right\rangle \end{split}$$

for every $f_1, \ldots, f_N, f'_1, \ldots, f'_{N'} \in L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3)$ and $N, N' \in \mathbb{N}$. From (3.3), it follows that the subspace spanned by the set of vectors $\left\{ \left[\prod_{\ell=1}^N b_{i_\ell j_\ell} (f_\ell)^* \right] \Psi_0 \right\}$ is dense in $\mathfrak{F}(L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3))$ and the subspace spanned by the set of vectors $\left\{ \left[\prod_{\ell=1}^N c_{i_\ell j_\ell} (f_\ell)^* \right] \bigotimes_{\lambda=1}^4 \Psi_\lambda \right\}$ is dense in $\bigotimes_{\lambda=1}^4 \mathfrak{H}_\lambda$. Hence, V can be extended to the unitary operator. To check (3.5) is easy.

Lemma 3.4. Let F be a real-valued even function on \mathbb{R}^3 . Assume that F is continuous. Then we obtain

$$V \mathrm{d}\Gamma(F \oplus F) V^{-1} = \sum_{\lambda=1}^{4} \mathrm{d}\Gamma_{\lambda}(F).$$

Proof. For readers' convenience, we will provide a sketch of the proof. We will continue to use the notations in the proof of Lemma 3.3. Set

$$egin{aligned} m{B}_{m{i}m{j}}(m{f}_1,\ldots,m{f}_N) &= \left[\prod_{\ell=1}^N b_{i_\ell j_\ell}(m{f}_\ell)^*
ight]\Psi_0, \ m{C}_{m{i}m{j}}(m{f}_1,\ldots,m{f}_N) &= \left[\prod_{\ell=1}^N c_{i_\ell j_\ell}(m{f}_\ell)^*
ight] \bigotimes_{\lambda=1}^4 \Psi_\lambda \end{aligned}$$

for $f_1, \ldots, f_N \in L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3)$, $i = (i_1, \ldots, i_N)$, $j = (j_1, \ldots, j_N) \in \{1, 2, 3\}^N$ and $N \in \mathbb{N}$. We define dense subspaces of $\mathfrak{F}(L^2(\mathbb{R}^3) \oplus L^2(\mathbb{R}^3))$ and $\bigotimes_{\lambda=1}^4 \mathfrak{H}_{\lambda}$ by

$$\mathfrak{V}_1 = \operatorname{Lin} \left\{ \boldsymbol{B}_{\boldsymbol{ij}}(\boldsymbol{f}_1, \dots, \boldsymbol{f}_N) \, \middle| \, \boldsymbol{f}_1, \dots, \boldsymbol{f}_N \in C_0(\mathbb{R}^3) \oplus C_0(\mathbb{R}^3), \, \boldsymbol{i}, \boldsymbol{j} \in \{1, 2, 3\}^N, \, N \in \mathbb{N} \right\}, \\ \mathfrak{V}_2 = \operatorname{Lin} \left\{ \boldsymbol{C}_{\boldsymbol{ij}}(\boldsymbol{f}_1, \dots, \boldsymbol{f}_N) \, \middle| \, \boldsymbol{f}_1, \dots, \boldsymbol{f}_N \in C_0(\mathbb{R}^3) \oplus C_0(\mathbb{R}^3), \, \boldsymbol{i}, \boldsymbol{j} \in \{1, 2, 3\}^N, \, N \in \mathbb{N} \right\}, (3.6)$$

where $\operatorname{Lin}(S)$ indicates the linear span of S. As is well-known, $\mathrm{d}\Gamma(F \oplus F)$ and $\sum_{\lambda=1}^{4} \mathrm{d}\Gamma_{\lambda}(F)$ are essentially self-adjoint on \mathfrak{V}_{1} and \mathfrak{V}_{2} , respectively. We readily confirm that

$$d\Gamma(F \oplus F) \boldsymbol{B}_{ij}(\boldsymbol{f}_1, \dots, \boldsymbol{f}_N) = \sum_{\alpha=1}^N \boldsymbol{B}_{ij}(\boldsymbol{f}_1, \dots, F \oplus F \boldsymbol{f}_\alpha, \dots, \boldsymbol{f}_N),$$
$$\sum_{\lambda=1}^4 d\Gamma_\lambda(F) \boldsymbol{C}_{ij}(\boldsymbol{f}_1, \dots, \boldsymbol{f}_N) = \sum_{\alpha=1}^N \boldsymbol{C}_{ij}(\boldsymbol{f}_1, \dots, F \oplus F \boldsymbol{f}_\alpha, \dots, \boldsymbol{f}_N).$$

Therefore, by Lemma 3.3, we obtain

$$V \mathrm{d}\Gamma(F \oplus F) V^{-1} = \sum_{\lambda=1}^{4} \mathrm{d}\Gamma_{\lambda}(F)$$
 on \mathfrak{V}_2 .

This concludes the proof of Lemma 3.4.

3.5 Definition of the Feynman Hamiltonians

In this subsection, we introduce the Feynman Hamiltonians. To this end, let

$$\begin{aligned} \mathscr{A}(x) &= \int_{\mathbb{R}^3} \mathrm{d}k \, \frac{\hat{\varrho}(k)}{\sqrt{2|k|}} \big\{ \varepsilon(k,1) \big(a(k,1) \cos(k \cdot x) + a(k,3) \sin(k \cdot x) \big) \\ &+ \varepsilon(k,2) \big(a(k,2) \cos(k \cdot x) + a(k,4) \sin(k \cdot x) \big) + \mathrm{h.c.} \big\}, \\ H_0 &= \sum_{\lambda=1}^4 \int_{\mathbb{R}^3} \mathrm{d}k \, |k| a(k,\lambda)^* a(k,\lambda). \end{aligned}$$

Here, h.c. denotes the hermite conjugates of the preceeding terms. Note that $\mathscr{A}(x)$ is essentially self-adjoint on \mathfrak{V}_2 defined by (3.6). We denote its closure by the same symbol. By (3.4) and (3.5), we have the following:

$$VA(x)V^{-1} = \mathscr{A}(x), \qquad VH_{\rm f}V^{-1} = H_0.$$

Now we define the two-electron Feynman Hamiltonian $H_{\rm F2e}$ by

$$H_{F2e} = \frac{1}{2} \left(-i\nabla_1 - e\mathscr{A}(0) \right)^2 + \frac{1}{2} e^2 \nu_0^2 x_1^2 + \frac{1}{2} \left(-i\nabla_2 - e\mathscr{A}(r) \right)^2 + \frac{1}{2} e^2 \nu_0^2 x_2^2 + e^2 \int_{\mathbb{R}^3} dk \, \hat{\varrho}(k)^2 e^{ik \cdot r} \left(x_1 \cdot \hat{k} \right) \left(x_2 \cdot \hat{k} \right) + H_0.$$

Remark that H_{F2e} acts in $L^2(\mathbb{R}^3_{x_1}) \otimes L^2(\mathbb{R}^3_{x_2}) \otimes \left(\bigotimes_{\lambda=1}^4 \mathfrak{F}(\mathfrak{H}_{\lambda})\right)$ and is bounded from below, provided that R is sufficiently large.

The following proposition plays an important role in the present paper.

Proposition 3.5. If R is large enough, $VH_{D2e}V^{-1} = H_{F2e}$.

Proof. Apply Lemmas 3.3 and 3.4.

As for the one-electron Feynman Hamiltonian, we obtain the following.

Proposition 3.6. Let

$$H_{\rm F1e} = \frac{1}{2} \left(-i\nabla - e\mathscr{A}(0) \right)^2 + \frac{1}{2} e^2 \nu_0^2 x^2 + H_0.$$

We have $VH_{\text{D1e}}V^{-1} = H_{\text{F1e}}$.

In Remark 7.2, we will explain why the Feynman Hamiltonians are useful.

4 Canonical transformations

Let U be a unitary operator on $L^2(\mathbb{R}^3_{x_1}) \otimes L^2(\mathbb{R}^3_{x_2}) \otimes \left(\bigotimes_{\lambda=1}^4 \mathfrak{F}(\mathfrak{H}_{\lambda})\right)$ defined by

$$U = \exp\{\mathrm{i}ex_1 \cdot \mathscr{A}(0) + \mathrm{i}ex_2 \cdot \mathscr{A}(r)\}.$$

Then one readily confirms that

$$U^*(-i\nabla_1)U = -i\nabla_1 + e\mathscr{A}(0), \qquad U^*(-i\nabla_2)U = -i\nabla_2 + e\mathscr{A}(r)$$

and

$$U^*a(k,\lambda)U = \begin{cases} a(k,\lambda) + ie\frac{\hat{\varrho}(k)}{\sqrt{2|k|}}\varepsilon(k,\lambda)\cdot(x_1 + x_2\cos(k\cdot r)) & \text{for } \lambda = 1,2, \\ a(k,\lambda) + ie\frac{\hat{\varrho}(k)}{\sqrt{2|k|}}\varepsilon(k,\lambda-2)\cdot x_2\sin(k\cdot r) & \text{for } \lambda = 3,4. \end{cases}$$

Here, we used the following fact:

$$e^T a(k, \lambda) e^{-T} = a(k, \lambda) + G(k, \lambda),$$

where $T = \sum_{\lambda=1}^{4} \{a_{\lambda}(G(\cdot,\lambda)) - a(G(\cdot,\lambda))^*\}^{**}, G(\cdot,\lambda) \in \mathfrak{H}_{\lambda}$. Hence, we arrive at⁴

$$U^{*}H_{F2e}U = -\frac{1}{2}\Delta_{1} + \frac{1}{2}e^{2}\nu^{2}x_{1}^{2} - \frac{1}{2}\Delta_{2} + \frac{1}{2}e^{2}\nu^{2}x_{2}^{2} + ex_{1} \cdot E(0) + ex_{2} \cdot E(r) + H_{0} + e^{2}\int_{\mathbb{R}^{3}} dk \,\hat{\varrho}(k)^{2}\cos(k \cdot r)(x_{1} \cdot x_{2}), \qquad (4.2)$$

where $\nu^2 = 2\nu_0^2$ and

$$E(x) = i \int_{\mathbb{R}^3} dk \sqrt{\frac{|k|}{2}} \hat{\varrho}(k) \{ \varepsilon(k,1) \big(\cos(k \cdot x)a(k,1) + \sin(k \cdot x)a(k,3) \big) + \varepsilon(k,2) \big(\cos(k \cdot x)a(k,2) + \sin(k \cdot x)a(k,4) \big) - \text{h.c.} \}.$$

Let $N_{\rm f}$ be the number operator defined by $N_{\rm f} = \sum_{\lambda=1}^{4} d\Gamma_{\lambda}(1)$. Applying the "Fourier transformation" $e^{-i\pi N_{\rm f}/2}$ in the Fock space,⁵ we obtain that

$$H = e^{i\pi N_{\rm f}/2} U^* H_{\rm F2e} U e^{-i\pi N_{\rm f}/2}$$

= $-\frac{1}{2} \Delta_1 + \frac{1}{2} e^2 \nu^2 x_1^2 - \frac{1}{2} \Delta_2 + \frac{1}{2} e^2 \nu^2 x_2^2 + ex_1 \cdot \hat{E}(0) + ex_2 \cdot \hat{E}(r) + H_0$
+ $e^2 \int_{\mathbb{R}^3} \mathrm{d}k \, \hat{\varrho}(k)^2 \cos(k \cdot r) (x_1 \cdot x_2),$ (4.3)

where

$$\hat{E}(x) = \int_{\mathbb{R}^3} \mathrm{d}k \, |k| \hat{\varrho}(k) \left\{ \varepsilon(k,1) \left(\cos(k \cdot x) q(k,1) + \sin(k \cdot x) q(k,3) \right) \right. \\ \left. + \varepsilon(k,2) \left(\cos(k \cdot x) q(k,2) + \sin(k \cdot x) q(k,4) \right) \right\}$$

$$(4.4)$$

and

$$q(k,\lambda) = \frac{1}{\sqrt{2|k|}} \left(a(k,\lambda) + a(k,\lambda)^* \right)$$

Since due to the assumption (A.3) the last term in (4.3) gives a rapidly decreasing contribution as a function of R to the ground state energy, we ignore this term from now on.

Finally, we define

$$K = -\frac{1}{2}\Delta + \frac{1}{2}e^{2}\nu^{2}x^{2} + ex \cdot \hat{E}(0) + H_{0}.$$

By an argument similar to the construction of U, we can construct a unitary operator u on $L^2(\mathbb{R}^3) \otimes \left(\bigotimes_{\lambda=1}^4 \mathfrak{F}(\mathfrak{H}_{\lambda})\right)$ such that $K = e^{i\pi N_f/2} u H_{F1e} u^{-1} e^{-i\pi N_f/2}$.

$$e^{2} \int_{\mathbb{R}^{3}} \mathrm{d}k \,\hat{\varrho}(k)^{2} \cos(k \cdot r) \bigg\{ \big(x_{1} \cdot \hat{k} \big) \big(x_{2} \cdot \hat{k} \big) + \sum_{\lambda=1,2} (\varepsilon(k,\lambda) \cdot x_{1}) (\varepsilon(k,\lambda) \cdot x_{2}) \bigg\}.$$

$$\tag{4.1}$$

Here, we used the fact that $\int dk \,\hat{\varrho}(k)^2 \sin(k \cdot r) (x_1 \cdot \hat{k}) (x_2 \cdot \hat{k}) = 0$. By applying the basic property $\sum_{\lambda=1,2} |\varepsilon(k,\lambda)\rangle \langle \varepsilon(k,\lambda)| = \mathbb{1}_3 - |\hat{k}\rangle \langle \hat{k}|$, we conclude that (4.1) is equal to $e^2 \int_{\mathbb{R}^3} dk \,\hat{\varrho}(k)^2 \cos(k \cdot r) (x_1 \cdot x_2)$.

 $\sum_{k=1}^{\lambda=1} \sum_{j=1}^{2} (a(k,\lambda) = -\frac{i}{\sqrt{2}} (a(k,\lambda) - a(k,\lambda)^*) \text{ and } \phi(k,\lambda) = \frac{1}{\sqrt{2}} (a(k,\lambda) + a(k,\lambda)^*).$ We can confirm that $[\pi(k,\lambda),\phi(k',\lambda')] = -i\delta_{\lambda\lambda'}\delta(k-k').$ Recalling the fact $[-id/dx,x] = -i, \pi(k,\lambda)$ and $\phi(k,\lambda)$ can be regarded as a multiplication operator and a differential operator, respectively. Now, we readily check that $e^{i\pi N_f/2}\pi(k,\lambda)e^{-i\pi N_f/2} = \phi(k,\lambda)$ holds, which corresponds to the relation $\mathcal{F}x\mathcal{F}^{-1} = -id/dx$, where \mathcal{F} is the Fourier transformation on $L^2(\mathbb{R})$. This similarity is a reason why we refer to the unitary operator $e^{i\pi N_f/2}$ as the Fourier transformation.

⁴The reason why the last term in the right-hand side of (4.2) appears is as follows. After performing the unitary transformation, we see that $U^*H_{\rm F2e}U$ contains the term concerning $(x_1\cdot\hat{k})(x_2\cdot\hat{k})$ and $(\varepsilon(k,\lambda)\cdot x_1)(\varepsilon(k,\lambda)\cdot x_2)$, which is given by

5 Lattice approximated Hamiltonians

In order to exactly compute the ground state energies of H and K, we will first introduce the lattice approximation of Hamiltonians. As we will see in later sections, the approximated Hamiltonians can be regarded as Hamiltonians of finite dimensional harmonic oscillator, which are exactly solvable.

For each $\Lambda > 0$, let χ_{Λ} be an ultraviolet cutoff function given by $\chi_{\Lambda}(k) = 1$ if $|k| \leq \Lambda$, $\chi_{\Lambda}(k) = 0$ otherwise. We define a linear operator $\hat{E}_{\Lambda}(x)$ by replacing $\hat{\varrho}(k)$ with $\hat{\varrho}(k)\chi_{\Lambda}(k)$ in the definition of $\hat{E}(x)$, i.e., the equation (4.4). We also define $H_{0,\Lambda}$ by

$$H_{0,\Lambda} = \sum_{\lambda=1}^{4} \int_{\mathbb{R}^3} \mathrm{d}k \, |k| \chi_{\Lambda}(k) a(k,\lambda)^* a(k,\lambda).$$

The Hamiltonians with a cutoff Λ are defined by

$$H_{\Lambda} = -\frac{1}{2}\Delta_{1} + \frac{1}{2}e^{2}\nu^{2}x_{1}^{2} - \frac{1}{2}\Delta_{2} + \frac{1}{2}e^{2}\nu^{2}x_{2}^{2} + ex_{1}\cdot\hat{E}_{\Lambda}(0) + ex_{2}\cdot\hat{E}_{\Lambda}(r) + H_{0,\Lambda},$$

$$K_{\Lambda} = -\frac{1}{2}\Delta + \frac{1}{2}e^{2}\nu^{2}x^{2} + ex\cdot\hat{E}_{\Lambda}(0) + H_{0,\Lambda}.$$

We readily see that H_{Λ} and K_{Λ} respectively converge to H and K in the norm resolvent sense as $\Lambda \to \infty$.

Let M be the (momentum) lattice with a cutoff Λ , namely,

$$M = \left\{ l \in (2\pi \mathbb{Z}/L)^3 \, \big| \, |l_i| \le 2\pi \Lambda, \, i = 1, 2, 3 \right\} \setminus \{0\}.$$

For later use, we label the elements of M as

$$M = \{k_1, \ldots, k_N\}.$$

Then the lattice approximated Hamiltonians are defined by

$$\begin{split} H_{L,\Lambda} &= -\frac{1}{2}\Delta_1 + \frac{1}{2}e^2\nu^2 x_1^2 - \frac{1}{2}\Delta_2 + \frac{1}{2}e^2\nu^2 x_2^2 + ex_1 \cdot \hat{E}_{L,\Lambda}(0) + ex_2 \cdot \hat{E}_{L,\Lambda}(r) + H_{0,L,\Lambda}, \\ K_{L,\Lambda} &= -\frac{1}{2}\Delta + \frac{1}{2}e^2\nu^2 x^2 + ex \cdot \hat{E}_{L,\Lambda}(0) + H_{0,L,\Lambda}, \end{split}$$

where

$$\hat{E}_{L,\Lambda}(x) = \left(\frac{2\pi}{L}\right)^{3/2} \sum_{k \in M} |k|\hat{\varrho}(k) \{\varepsilon(k,1) (\cos(k \cdot x)q(k,1) + \sin(k \cdot x)q(k,3)) + \varepsilon(k,2) (\cos(k \cdot x)q(k,2) + \sin(k \cdot x)q(k,4))\},$$
(5.1)
$$H_{0,L,\Lambda} = \frac{1}{2} \sum_{\lambda=1}^{4} \sum_{k \in M} \left(p(k,\lambda)^2 + |k|^2 q(k,\lambda)^2\right) - 2 \sum_{k \in M} |k|$$

with $p(k,\lambda) = \frac{1}{i}\sqrt{\frac{|k|}{2}}(a(k,\lambda) - a(k,\lambda)^*)$. The lattice approximated operators act in the Hilbert space $L^2(\mathbb{R}^3_{x_1}) \otimes L^2(\mathbb{R}^3_{x_2}) \otimes \left(\bigotimes_{\lambda=1}^4 \mathfrak{F}(\mathfrak{H}_{L,\Lambda,\lambda})\right)$ or $L^2(\mathbb{R}^3) \otimes \left(\bigotimes_{\lambda=1}^4 \mathfrak{F}(\mathfrak{H}_{L,\Lambda,\lambda})\right)$, where $\mathfrak{H}_{L,\Lambda,\lambda} = \ell^2_*(M)$ $\cap \mathfrak{H}_{\lambda}$. Here, $\ell^2_*(M)$ is the $\ell^2(M)$ equipped with a modified norm

$$||f||_* = \left(\frac{2\pi}{L}\right)^{3/4} \left(\sum_{k \in M} |f(k)|^2\right)^{1/2},$$

and we regard $\ell^2_*(M)$ as a closed subspace of $L^2(\mathbb{R}^3)$. Note that $p(k, \lambda)$ and $q(k, \lambda)$ are essentially self-adjoint on the finite particle subspace of $\bigotimes_{\lambda=1}^{4} \mathfrak{F}(\mathfrak{H}_{L,\Lambda,\lambda})$. In what follows, we denote their closures by same symbols, respectively. $q(k, \lambda)$ and $p(k, \lambda)$ is a canonical pair of the photonic displacement coordinate and its conjugate momentum satisfying the standard commutation relations:

$$\begin{split} &[p(k,\lambda),q(k',\lambda')] = -\mathrm{i}\delta_{kk'}\delta_{\lambda\lambda'}, \\ &[p(k,\lambda),p(k',\lambda')] = 0 = [q(k,\lambda),q(k',\lambda')]. \end{split}$$

Recall the identification $\mathfrak{F}(\mathbb{C}) = L^2(\mathbb{R})$. Using this, we can naturally embed $\bigotimes_{\lambda=1}^4 \mathfrak{F}(\mathfrak{H}_{L,\Lambda,\lambda})$ into

 $\left(\bigotimes_{\lambda=1}^{4} L^2(\mathbb{R}^3)\right)^{\otimes \#M}$. In addition, $p(k,\lambda)$ and $q(k,\lambda)$ can be regarded as the differential and multiplication operators, respectively.

The following proposition is a basis for our computation.

Proposition 5.1. For each $z \in \mathbb{C} \setminus \mathbb{R}$, one has

$$\lim_{\Lambda \to \infty} \lim_{L \to \infty} (H_{L,\Lambda} - z)^{-1} = (H - z)^{-1},$$
$$\lim_{\Lambda \to \infty} \lim_{L \to \infty} (K_{L,\Lambda} - z)^{-1} = (K - z)^{-1}$$

in the operator norm topology.

Proof. See, e.g., [3, 9].

6 Diagonalization I: One-electron Hamiltonian

In this section, we diagonalize the one-electron Hamiltonian $K_{L,\Lambda}$. To this end, let

$$F_x(k,1) = \left(\frac{2\pi}{L}\right)^{3/2} |k|\hat{\varrho}(k)\cos(k\cdot x), \qquad F_x(k,2) = \left(\frac{2\pi}{L}\right)^{3/2} |k|\hat{\varrho}(k)\cos(k\cdot x),$$
$$F_x(k,3) = \left(\frac{2\pi}{L}\right)^{3/2} |k|\hat{\varrho}(k)\sin(k\cdot x), \qquad F_x(k,4) = \left(\frac{2\pi}{L}\right)^{3/2} |k|\hat{\varrho}(k)\sin(k\cdot x).$$

We define a linear operator $\mathbb{T}(x)$ from $\ell^2(M \times \{1, \ldots, 4\})$ to \mathbb{C}^3 by

$$\mathbb{T}(x)\boldsymbol{f} = \sum_{\lambda=1}^{4} \sum_{k \in M} |\varepsilon(k,\lambda)\rangle F_x(k,\lambda) f(k,\lambda)$$

for each $\mathbf{f} = \{f(k,\lambda) \mid k \in M, \lambda \in \{1, \dots, 4\}\} \in \ell^2(M \times \{1, \dots, 4\})$. Here, we used the following notation: $\varepsilon(k,3) := \varepsilon(k,1)$ and $\varepsilon(k,4) := \varepsilon(k,2)$. The adjoint of $\mathbb{T}(x)$ is denoted by $\mathbb{T}^*(x)$. Note that

$$(\mathbb{T}^*(x)a)(k,\lambda) = \langle \varepsilon(k,\lambda) | a \rangle_3 F_x(k,\lambda), \qquad a \in \mathbb{C}^3,$$

where $\langle \cdot | \cdot \rangle_3$ stands for the inner product in \mathbb{C}^3 .

Using the above notations, the interaction term $x \cdot \hat{E}_{L,\Lambda}(r)$ in $K_{L,\Lambda}$ is expressed as $x \cdot \hat{E}_{L,\Lambda}(r) = \langle \mathbb{T}(r)\boldsymbol{q} | x \rangle_3 = \langle x | \mathbb{T}(r)\boldsymbol{q} \rangle_3$, where $\boldsymbol{q} = \{q(k,\lambda) | k \in M, \lambda \in \{1,\ldots,4\}\}$. On the other hand, the field energy can be represented by

$$H_{0,L,\Lambda} = rac{1}{2} ig(oldsymbol{p}^2 + \langle oldsymbol{q} | S_0 oldsymbol{q}
angle ig) - rac{1}{2} \operatorname{tr}ig[\sqrt{S_0} ig],$$

where $p = \{p(k, \lambda) | k \in M, \lambda \in \{1, ..., 4\}\}$ and

$$S_0 = \begin{pmatrix} |k_1|^2 \mathbb{1}_4 & & O \\ & |k_2|^2 \mathbb{1}_4 & & \\ & & \ddots & \\ & & & \ddots & \\ & & & & |k_N|^2 \mathbb{1}_4 \end{pmatrix}.$$

Hence, $K_{L,\Lambda}$ can be rewritten as

$$K_{L,\Lambda} = -\frac{1}{2}\Delta + \frac{1}{2}e^2\nu^2 x^2 + e\langle x | \mathbb{T}(r)\boldsymbol{q}\rangle_3 + \frac{1}{2}\left(\boldsymbol{p}^2 + \langle \boldsymbol{q} | S_0\boldsymbol{q}\rangle\right) - \frac{1}{2}\operatorname{tr}\left[\sqrt{S_0}\right].$$
(6.1)

By setting $\phi = (x, q)$ and $\pi = (-i\nabla, p)$, one sees that

$$K_{L,\Lambda} = \frac{1}{2} \left(\langle \pi | \pi \rangle + \langle \phi | \omega \phi \rangle \right) - \frac{1}{2} \operatorname{tr} \left[\sqrt{\omega_0} \right] + \frac{3}{2} e \nu$$

where

$$\omega = \omega_0 + Q, \qquad \omega_0 = \begin{pmatrix} e^2 \nu^2 & 0\\ 0 & S_0 \end{pmatrix}, \qquad Q = e \begin{pmatrix} 0 & \mathbb{T}(r)\\ \mathbb{T}^*(r) & 0 \end{pmatrix}.$$

The following lemma is a basic input.

Lemma 6.1. If $1 \le \sqrt{2}e\nu_0$ and $\sqrt{2}e\|\hat{\varrho}\|_* < 1$, then $\omega \ge 0$.

Proof. By (A.3), we have $\|\mathbb{T}(r)\boldsymbol{f}\| \leq \sqrt{2}\|\hat{\varrho}\|_* \|S_0^{1/2}\boldsymbol{f}\|$ for all $\boldsymbol{f} \in \ell^2(M \times \{1,\ldots,4\})$. Hence, for all $\varphi = (a, \boldsymbol{f}) \in \mathbb{C}^3 \oplus \ell^2(M \times \{1,\ldots,4\})$, we have, by the Schwarz inequality,

$$\begin{split} |\langle \varphi | Q\varphi \rangle| &\leq 2\sqrt{2}e \|\hat{\varrho}\|_* \|a\|_3 \left\| S_0^{1/2} \boldsymbol{f} \right\| \\ &\leq \sqrt{2}e \|\hat{\varrho}\|_* \left(\|a\|_3^2 + \left\| S_0^{1/2} \boldsymbol{f} \right\|^2 \right) \\ &\leq \sqrt{2}e \|\hat{\varrho}\|_* \langle \varphi | \omega_0 \varphi \rangle, \end{split}$$

provided that $1 \le e^2 \nu^2$. This concludes the proof of Lemma 6.1.

Therefore, the ground state energy of $K_{L,\Lambda}$ is given by the following formula.

Proposition 6.2. Let $E_{L,\Lambda} = \inf \operatorname{spec}(K_{L,\Lambda})$. If $1 \leq \sqrt{2}e\nu_0$ and $\sqrt{2}e\|\hat{\varrho}\|_* < 1$, then one has

$$E_{L,\Lambda} = \frac{1}{2} \operatorname{tr} \left[\sqrt{\omega} - \sqrt{\omega_0} \right] + \frac{3}{2} e\nu$$

Proof. We provide a sketch of the proof. First, we diagonalize ω as

$$\omega = U^{-1} \operatorname{diag}(\lambda_1, \dots, \lambda_{4N+3}) U,$$

where U is a unitary matrix and $\lambda_1, \ldots, \lambda_{4N+3}$ are positive eigenvalues of ω . By setting $\tilde{\phi} = U\phi$ and $\tilde{\pi} = U\pi$, we can express $K_{L,\Lambda}$ as

$$K_{L,\Lambda} = \frac{1}{2} \left\langle \tilde{\pi} | \tilde{\pi} \right\rangle + \frac{1}{2} \left\langle \tilde{\phi} | \operatorname{diag}(\lambda_1, \dots, \lambda_{4N+3}) \tilde{\phi} \right\rangle - \frac{1}{2} \operatorname{tr}[\sqrt{\omega_0}] + \frac{3}{2} e\nu.$$
(6.2)

Because $\tilde{\pi}_j$ and $\tilde{\phi}_j$ satisfy the Weyl relation: $e^{it\tilde{\pi}_i}e^{is\tilde{\phi}_j} = e^{ist\delta_{ij}}e^{is\tilde{\phi}_j}e^{it\tilde{\pi}_i}$, the von Neumann's uniqueness theorem [23, Theorem VIII.14] tells us that there is a unitary operator $\tau \colon L^2(\mathbb{R}^{4N+3}) \to L^2(\mathbb{R}^{4N+3})$ such that $\tau\tilde{\phi}_j\tau^{-1} = x_j$ and $\tau\tilde{\pi}_j\tau^{-1} = -i\partial/\partial x_j$. Therefore, the right-hand side

of (6.2) can be regarded as a Hamiltonian for 4N + 3-dimensional harmonic oscillator. Since the lowest eigenvalue of the Hamiltonian $-\frac{1}{2}\Delta_j^2 + \frac{\lambda_j}{2}x_j^2$ is equal to $\sqrt{\lambda_j}/2$, we obtain that

$$E_{L,\Lambda} = \frac{1}{2} \sum_{j=1}^{4N+3} \sqrt{\lambda_j} - \frac{1}{2} \operatorname{tr} \left[\sqrt{\omega_0} \right] + \frac{3}{2} e\nu = \frac{1}{2} \operatorname{tr} \left[\sqrt{\omega} - \sqrt{\omega_0} \right] + \frac{3}{2} e\nu.$$

This finishes the proof of Proposition 6.2.

Applying the elementary fact

$$\frac{1}{\pi} \int_{-\infty}^{\infty} \mathrm{d}s \, \frac{a}{s^2 + a} = \sqrt{a},\tag{6.3}$$

we have that

$$E_{L,\Lambda} = \frac{1}{2\pi} \int_{-\infty}^{\infty} \mathrm{d}s \, \mathrm{tr} \left[\omega \left(s^2 + \omega \right)^{-1} - \omega_0 \left(s^2 + \omega_0 \right)^{-1} \right] + \frac{3}{2} e\nu$$

$$= \frac{1}{2\pi} \sum_{j=1}^{\infty} (-1)^j \int_{-\infty}^{\infty} \mathrm{d}s \, \mathrm{tr} \left[\omega_0 \left(s^2 + \omega_0 \right)^{-1} \left\{ Q \left(s^2 + \omega_0 \right)^{-1} \right\}^j \right]$$

$$+ \frac{1}{2\pi} \sum_{j=1}^{\infty} (-1)^{j+1} \int_{-\infty}^{\infty} \mathrm{d}s \, \mathrm{tr} \left[\left\{ Q \left(s^2 + \omega_0 \right)^{-1} \right\}^j \right] + \frac{3}{2} e\nu$$

$$= \frac{1}{2\pi} \sum_{j=1}^{\infty} (-1)^{j+1} \int_{-\infty}^{\infty} \mathrm{d}s \, s^2 \, \mathrm{tr} \left[\left(s^2 + \omega_0 \right)^{-1} \left\{ Q \left(s^2 + \omega_0 \right)^{-1} \right\}^j \right] + \frac{3}{2} e\nu.$$
(6.4)

Since Q is off-diagonal, (6.4) becomes

$$E_{L,\Lambda} = -\frac{1}{2\pi} \sum_{j=1}^{\infty} \int_{-\infty}^{\infty} \mathrm{d}s \, s^2 \, \mathrm{tr} \left[\left(s^2 + \omega_0 \right)^{-1} Q(s)^{2j} \right] + \frac{3}{2} e\nu, \tag{6.5}$$

where $Q(s) = (s^2 + \omega_0)^{-1/2} Q(s^2 + \omega_0)^{-1/2}$. In what follows, we will examine the convergence of the right-hand side of (6.5). As we will see, this series absolutely converges and (6.5) is rigorously justified if ν_0 is large enough.

We begin with the following basic lemma.

Lemma 6.3. We have the following

$$\|\mathbb{T}(x)(s^2+S_0)^{-1/2}\| \le \sqrt{2}\|\hat{\varrho}\|_*, \qquad \|(s^2+S_0)^{-1/2}\mathbb{T}^*(x)\| \le \sqrt{2}\|\hat{\varrho}\|_*$$

for all $x \in \mathbb{R}^3$, where $||f||_* = \sqrt{\left(\frac{2\pi}{L}\right)^3 \sum_{k \in M} |f(k)|^2}$ for each $f \in \ell^2(M)$.

Proof. For each $f \in \ell^2(M \times \{1, \dots, 4\})$, we have, by (A.3),

$$\begin{split} \left\| \mathbb{T}(x) \left(s^{2} + S_{0} \right)^{-1/2} \boldsymbol{f} \right\|^{2} &= \langle \boldsymbol{f} \mid \mathbb{T}_{s}^{*}(x) \mathbb{T}_{s}(x) \boldsymbol{f} \rangle \\ &= \langle \boldsymbol{f} \mid \left(s^{2} + S_{0} \right)^{-1/2} \mathbb{M}(x, x) \left(s^{2} + S_{0} \right)^{-1/2} \boldsymbol{f} \rangle \\ &\leq \left| \sum_{k,\lambda} \frac{F_{x}(k,\lambda)}{\left(s^{2} + k^{2} \right)^{1/2}} \boldsymbol{f}(k,\lambda) \right|^{2} \\ &\leq \left\| \left(s^{2} + k^{2} \right)^{-1/2} F_{x} \right\|^{2} \|\boldsymbol{f}\|^{2}. \end{split}$$

Because $\|(s^2 + k^2)^{-1/2} F_x\|^2 \le 2\|\hat{\varrho}\|_*^2$, we conclude that $\|\mathbb{T}(x)(s^2 + S_0)^{-1/2}\| \le \sqrt{2}\|\hat{\varrho}\|_*$.

Lemma 6.4. Let

$$D(s) = 2e^{2}s^{2}(s^{2} + e^{2}\nu^{2})^{-1} \times \{(s^{2} + e^{2}\nu^{2})^{-1} \| (s^{2} + |k|^{2})^{-1/2} |k|\hat{\varrho}\|_{*}^{2} + \| (s^{2} + |k|^{2})^{-1} |k|\hat{\varrho}\|_{*}^{2} \}.$$
(6.6)

Then we have the following:

(i) For all $s \in \mathbb{R}$,

$$s^{2} \operatorname{tr} \left[\left(s^{2} + \omega_{0} \right)^{-1} Q(s)^{2n} \right] \leq \left(\frac{\sqrt{2}}{\nu} \| \hat{\varrho} \|_{*} \right)^{2n-2} D(s).$$

(ii) Let $a = \left(\frac{\sqrt{2}}{\nu} \|\hat{\varrho}\|_*\right)^2$. If a < 1, then we have

$$\sum_{n=1}^{\infty} s^2 \operatorname{tr} \left[\left(s^2 + \omega_0 \right)^{-1} Q(s)^{2n} \right] \le \frac{1}{1-a} D(s).$$

Remark that $\lim_{L\to\infty} a \leq \left(\frac{\sqrt{2}}{\nu} \|\hat{\varrho}\|_{L^2}\right)^2 \leq c_{\infty}^2 < 1/4$ holds for all $\Lambda > 0$ by the assumption in Theorem 2.1. Thus, the condition a < 1 is satisfied provided that L is sufficiently large.

(iii) $D(s) \in L^1(\mathbb{R})$ and

$$\frac{1}{2\pi} \int_{\mathbb{R}} \mathrm{d}s \, D(s) \le \frac{e}{\nu} \|\hat{\varrho}\|_*^2. \tag{6.7}$$

Proof. We set $\mathbb{T}_s(r) = \mathbb{T}(r)(s^2 + S_0)^{-1/2}$. First, consider the case where n = 1. Because

$$Q(s)^{2} = e^{2} \left(s^{2} + e^{2} \nu^{2}\right)^{-1} \begin{pmatrix} \mathbb{T}_{s}(r) \mathbb{T}_{s}^{*}(r) & 0\\ 0 & \mathbb{T}_{s}^{*}(r) \mathbb{T}_{s}(r) \end{pmatrix},$$
(6.8)

we obtain that

$$s^{2} \operatorname{tr} \left[\left(s^{2} + \omega_{0} \right)^{-1} Q(s)^{2} \right] = e^{2} s^{2} \left(s^{2} + e^{2} \nu^{2} \right)^{-2} \operatorname{tr} \left[\mathbb{T}_{s}(r) \mathbb{T}_{s}^{*}(r) \right] + e^{2} s^{2} \left(s^{2} + e^{2} \nu^{2} \right)^{-1} \operatorname{tr} \left[\left(s^{2} + S_{0} \right)^{-1} \mathbb{T}_{s}^{*}(r) \mathbb{T}_{s}(r) \right].$$

By (A.3) and (A.4), we have

$$\operatorname{tr}\left[\mathbb{T}_{s}(r)\mathbb{T}_{s}^{*}(r)\right] \leq 2\left(\frac{2\pi}{L}\right)^{3} \sum_{k \in M} \left(s^{2} + |k|^{2}\right)^{-1} |k|^{2} |\hat{\varrho}(k)|^{2}, \\ \operatorname{tr}\left[\left(s^{2} + S_{0}\right)^{-1} \mathbb{T}_{s}^{*}(r)\mathbb{T}_{s}(r)\right] \leq 2\left(\frac{2\pi}{L}\right)^{3} \sum_{k \in M} \left(s^{2} + |k|^{2}\right)^{-2} |k|^{2} |\hat{\varrho}(k)|^{2}$$

Thus, we get (i) for n = 1.

To prove the assertion for $n \ge 2$, we remark that $\|Q(s)\| \le \frac{\sqrt{2}}{\nu} \|\hat{\varrho}\|_*$, which immediately follows from Lemma 6.3 and (6.8). Thus, by using the fact $Q(s)^{2n} \le \|Q(s)\|^{2n-2}Q(s)^2$, we have

$$\operatorname{tr}\left[\left(s^{2}+\omega_{0}\right)^{-1/2}Q(s)^{2n}\left(s^{2}+\omega_{0}\right)^{-1/2}\right] \leq \|Q(s)\|^{2n-2}\operatorname{tr}\left[\left(s^{2}+\omega_{0}\right)^{-1/2}Q(s)^{2}\left(s^{2}+\omega_{0}\right)^{-1/2}\right].$$

Applying the result for n = 1, we get the desired result for $n \ge 2$. (ii) immediately follows from (i).

By using the formula (C.1) with $a = b = e^2 \nu^2$ and $c = |k|^2$, we see that

$$\frac{1}{2\pi} \int_{-\infty}^{\infty} \mathrm{d}s \, 2e^2 s^2 \left(s^2 + e^2 \nu^2\right)^{-2} \left\| \left(s^2 + |k|^2\right)^{-1/2} |k|\hat{\varrho}\|_*^2 \\ = \frac{e^2}{2\pi} \left(\frac{2\pi}{L}\right)^3 \sum_{k \in M} \frac{2\pi}{2e\nu} \frac{|k|^2 |\hat{\varrho}(k)|^2}{(e\nu + |k|)^2} \le \frac{e}{2\nu} \|\hat{\varrho}\|_*^2.$$

Similarly, by using the formula (C.1) with $a = e^2 \nu^2$ and $b = c = |k|^2$, we obtain

$$\frac{1}{2\pi} \int_{-\infty}^{\infty} \mathrm{d}s \, 2e^2 s^2 \left(s^2 + e^2 \nu^2\right)^{-1} \left\| \left(s^2 + |k|^2\right)^{-1} |k|\hat{\varrho}\|_*^2 \\ = \frac{e^2}{2\pi} \left(\frac{2\pi}{L}\right)^3 \sum_{k \in M} \frac{2\pi}{2|k|(|k| + e\nu)^2} |k|^2 |\hat{\varrho}(k)|^2 \le \frac{e}{2\nu} \|\hat{\varrho}\|_*^2.$$

Inserting these into (6.6), we obtain the assertion (iii).

Corollary 6.5. The right-hand side of (6.5) absolutely converges, provided that $\sqrt{2}\|\hat{\varrho}\|_* < \nu$, $1 \leq \sqrt{2}e\nu_0$ and $\sqrt{2}e\|\hat{\varrho}\|_* < 1$. In addition, to exchange the series with the integral in (6.5) (or (6.4)) can be justified.

7 Diagonalization II: Two-electron Hamiltonian

Next we will diagonalize $H_{L,\Lambda}$. This is actually possible because we employ the Feynman Hamiltonian, see Remark 7.2 for details. By an argument similar to that of the proof of (6.1), $H_{L,\Lambda}$ can be expressed as

$$H_{L,\Lambda} = -\frac{1}{2}\Delta_1 + \frac{1}{2}e^2\nu^2 x_1^2 - \frac{1}{2}\Delta_2 + \frac{1}{2}e^2\nu^2 x_2^2 + e\langle \mathbb{T}(0)\boldsymbol{q}|x_1\rangle_3 + e\langle \mathbb{T}(r)\boldsymbol{q}|x_2\rangle_3 + \frac{1}{2}(\boldsymbol{p}^2 + \langle \boldsymbol{q}|S_0\boldsymbol{q}\rangle) - \frac{1}{2}\operatorname{tr}\left[\sqrt{S_0}\right].$$

By setting $\Phi = (x_1, x_2, q)$ and $\Pi = (-i\nabla_1, -i\nabla_2, p)$, we have that

$$H_{L,\Lambda} = \frac{1}{2} \left(\langle \Pi | \Pi \rangle + \langle \Phi | \Omega \Phi \rangle \right) - \frac{1}{2} \operatorname{tr} \left[\sqrt{\Omega_0} \right] + 3e\nu,$$

where

$$\begin{split} \Omega &= \Omega_0 + Q_1 + Q_2, \\ \Omega_0 &= \begin{pmatrix} e^2 \nu^2 & 0 & 0 \\ 0 & e^2 \nu^2 & 0 \\ 0 & 0 & S_0 \end{pmatrix}, \qquad Q_1 = e \begin{pmatrix} 0 & 0 & \mathbb{T}(0) \\ 0 & 0 & 0 \\ \mathbb{T}^*(0) & 0 & 0 \end{pmatrix}, \qquad Q_2 = e \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & \mathbb{T}(r) \\ 0 & \mathbb{T}^*(r) & 0 \end{pmatrix}. \end{split}$$

By an argument similar to that in the proof of Proposition 6.2, we get the following useful formula.

Proposition 7.1. Let $E_{L,\Lambda}(R) = \inf \operatorname{spec}(H_{L,\Lambda})$. If $1 \leq \sqrt{2}e\nu_0$ and $\sqrt{2}e\|\hat{\varrho}\|_* < 1$, then $\Omega \geq 0$ and

$$E_{L,\Lambda}(R) = \frac{1}{2} \operatorname{tr} \left[\sqrt{\Omega} - \sqrt{\Omega_0} \right] + 3e\nu.$$

Remark 7.2 (Why are the Feynman Hamiltonians helpful?). From the expression (5.1), we see that $\hat{E}_{L,\Lambda}(x)$ can be written as a sum of multiplication operators $q(k,\lambda)$. As we already knew, this fact is a key to the diagonalization of $H_{L,\Lambda}$. In contrast to the Feynman Hamiltonians, in the standard representation, $\hat{E}_{L,\Lambda}(x)$ corresponds to the following operator:

$$\left(\frac{2\pi}{L}\right)^{3/2} \sum_{\lambda=1,2} \sum_{k \in M} \hat{\varrho}(k) \varepsilon(k,\lambda) \big\{ \cos(k \cdot x) |k| q(k,\lambda) + \sin(k \cdot x) |k|^{-1} p(k,\lambda) \big\}.$$
(7.1)

In (7.1), both multiplication and differential operators appear, provided that $x \neq 0$. At first glance, it appears that diagonalizing the Hamiltonians in this representation requires extra efforts.

Moreover, it can be readily seen that, by (6.3),

$$E_{L,\Lambda}(R) = \frac{1}{2\pi} \sum_{j=1}^{\infty} (-1)^{j+1} \int_{-\infty}^{\infty} \mathrm{d}s \, s^2 \, \mathrm{tr} \left[\left(s^2 + \Omega_0 \right)^{-1} \left\{ (Q_1 + Q_2) \left(s^2 + \Omega_0 \right)^{-1} \right\}^j \right] + 3e\nu.$$
(7.2)

To examine this formal series, let us introduce the following notation:

$$\langle O_1 O_2 \cdots O_n \rangle = \frac{1}{2\pi} \int_{-\infty}^{\infty} \mathrm{d}s \, s^2 \, \mathrm{tr} \left[\left(s^2 + \Omega_0 \right)^{-1} O_1(s) O_2(s) \cdots O_n(s) \right],$$

where $O_i(s) = (s^2 + \Omega_0)^{-1/2} O_i (s^2 + \Omega_0)^{-1/2}$. Then (7.2) can be expressed as

$$E_{L,\Lambda}(R) = \sum_{j=1}^{\infty} (-1)^{j+1} \langle \underbrace{(Q_1 + Q_2) \cdots (Q_1 + Q_2)}_{j} \rangle + 3e\nu$$

$$= \sum_{j=1}^{\infty} (-1)^{j+1} \langle \underbrace{Q_1 \cdots Q_1}_{j} \rangle + \sum_{j=1}^{\infty} (-1)^{j+1} \langle \underbrace{Q_2 \cdots Q_2}_{j} \rangle$$

$$+ \sum_{j=1}^{\infty} \sum_{\substack{i_1, \dots, i_j \in \{1, 2\}\\\{i_1, \dots, i_j\} \neq \{1, 1, \dots, 1\}, \{2, 2, \dots, 2\}}} (-1)^{j+1} \langle Q_{i_1} \cdots Q_{i_j} \rangle + 3e\nu.$$
(7.3)

Since Q_1 and Q_2 are off-diagonal, we have

$$E_{L,\Lambda}(R) = -\sum_{j=1}^{\infty} \langle \underbrace{Q_1 \cdots Q_1}_{2j} \rangle - \sum_{j=1}^{\infty} \langle \underbrace{Q_2 \cdots Q_2}_{2j} \rangle$$
$$-\sum_{j=1}^{\infty} \sum_{\substack{i_1, \dots, i_{2j} \in \{1, 2\}\\\{i_1, \dots, i_{2j}\} \neq \{1, 1, \dots, 1\}, \{2, 2, \dots, 2\}}} \langle Q_{i_1} \cdots Q_{i_{2j}} \rangle + 3e\nu.$$

On the other hand, we remark that, by Corollary 6.5,

$$E_{L,\Lambda} = -\sum_{j=1}^{\infty} \langle \underline{Q_1 \cdots Q_1}_{2j} \rangle + \frac{3}{2} e\nu = -\sum_{j=1}^{\infty} \langle \underline{Q_2 \cdots Q_2}_{2j} \rangle + \frac{3}{2} e\nu,$$

provided that $\sqrt{2}\|\hat{\varrho}\|_* < \nu$, $1 \leq \sqrt{2}e\nu_0$ and $\sqrt{2}e\|\hat{\varrho}\|_* < 1$. Thus, we formally arrive at the following formula:

$$E_{L,\Lambda}(R) - 2E_{L,\Lambda} = -\sum_{j=1}^{\infty} \sum_{\substack{i_1,\dots,i_{2j} \in \{1,2\}\\\{i_1,\dots,i_{2j}\} \neq \{1,1,\dots,1\},\{2,2,\dots,2\}}} \langle Q_{i_1} \cdots Q_{i_{2j}} \rangle.$$
(7.4)

Our next task is to prove the convergence of the right-hand side of (7.4). For this purpose, we need some preliminaries. Let

$$\mathcal{I}_{2j} = \{I = \{i_1, \dots, i_{2j}\}, i_1, \dots, i_{2j} \in \{1, 2\} \mid I \neq \{1, 1, \dots, 1\}, \{2, 2, \dots, 2\}\}.$$

For each $I \in \mathcal{I}_{2j}$, we set $|I| = i_1 + i_2 + \cdots + i_{2j}$. Furthermore, we use the following notation:

$$Q_I = Q_{i_1} Q_{i_2} \cdots Q_{i_{2j}}, \qquad I = \{i_1, \dots, i_{2j}\} \in \mathcal{I}_{2j}.$$

Lemma 7.3. Let $I \in \mathcal{I}_{2j}$. If |I| is an odd number, then $\langle Q_I \rangle = 0$.

Proof. Note that $(s^2 + \Omega_0)^{-1}$, $Q_1(s)^2$ and $Q_2(s)^2$ are diagonal operators, while $Q_1(s)Q_2(s)$ and $Q_2(s)Q_1(s)$ are off-diagonal operators, see Appendix A. Hence, if |I| is an odd number, then $(s^2 + \Omega_0)^{-1}Q_{i_1}(s)\cdots Q_{i_{2j}}(s)$ is an off-diagonal operator. Accordingly,

tr
$$[(s^2 + \Omega_0)^{-1}Q_{i_1}(s)\cdots Q_{i_{2j}}(s)] = 0.$$

This concludes the proof of Lemma 7.3.

Let $\mathcal{I}_{2j}^{(e)} = \{I \in \mathcal{I}_{2j} \mid |I| \text{ is even}\}$. By Lemma 7.3, we have

the r.h.s. of (7.4) =
$$-\sum_{j=1}^{\infty} \sum_{I \in \mathcal{I}_{2j}^{(e)}} \langle Q_I \rangle.$$
 (7.5)

Lemma 7.4. For each $s \in \mathbb{R}$ and $I = \{i_1, \ldots, i_{2j}\} \in \mathcal{I}_{2j}^{(e)}$, we set

$$Q_I(s) = Q_{i_1}(s) \cdots Q_{i_{2j}}(s)$$

and

$$E_{I}(s) = s^{2} \operatorname{tr} \left[\left(s^{2} + \omega \right)^{-1} Q_{I \setminus \{i_{1}\}}^{*}(s) Q_{I \setminus \{i_{1}\}}(s) \right]$$

where $Q^*_{I\setminus\{i_1\}}(s) = (Q_{I\setminus\{i_1\}}(s))^*$. For all R > 0, we have the following:

(i) For each $s \in \mathbb{R}$ and $I \in \mathcal{I}_{2j}^{(e)}$,

$$s^{2} | \operatorname{tr} \left[\left(s^{2} + \omega \right)^{-1} Q_{I}(s) \right] | \leq D(s)^{1/2} E_{I}(s)^{1/2},$$
(7.6)

where D(s) is given by (6.6).

(ii) Recall that a is defined by $a = \left(\frac{\sqrt{2}}{\nu} \|\hat{\varrho}\|_*\right)^2$. If a < 1/4, then

$$\sum_{j=1}^{\infty} \sum_{I \in \mathcal{I}_{2j}^{(e)}} E_I(s)^{1/2} \le D(s)^{1/2} \frac{4}{1-4a}.$$
(7.7)

Thus,
$$D(s)^{1/2} \sum_{j=1}^{\infty} \sum_{I \in \mathcal{I}_{2j}^{(e)}} E_I(s)^{1/2} \in L^1(\mathbb{R})$$
 and
$$\sum_{j=1}^{\infty} \sum_{I \in \mathcal{I}_{2j}^{(e)}} |\langle Q_I \rangle| \le \frac{e}{\nu} \|\hat{\varrho}\|_*^2 \frac{4}{1-4a}.$$

Note that as we mentioned in Lemma 6.4, the condition a < 1/4 is satisfied provided that L is large enough.

Proof. For notational simplicity, we set $G = (s^2 + \omega_0)^{-1}$. By the Schwarz inequality $|\operatorname{tr}[A^*B]| \leq \operatorname{tr}[A^*A]^{1/2} \operatorname{tr}[B^*B]^{1/2}$, we obtain

$$\operatorname{tr} \left[G^{1/2} Q_{i_1}(s) \cdots Q_{i_{2j}}(s) G^{1/2} \right] | \leq \operatorname{tr} \left[G^{1/2} Q_{i_1}(s) Q_{i_1}(s) G^{1/2} \right]^{1/2} \\ \times \operatorname{tr} \left[G^{1/2} Q_{i_{2j}}(s) Q_{i_{2j-1}}(s) \cdots Q_{i_2}(s) Q_{i_2}(s) \cdots Q_{i_{2j}}(s) G^{1/2} \right]^{1/2},$$

which implies that

$$s^{2} |\operatorname{tr} \left[\left(s^{2} + \omega \right)^{-1} Q_{I}(s) \right] | \leq \left\{ s^{2} \operatorname{tr} \left[\left(s^{2} + \omega \right)^{-1} Q_{i_{1}}(s) Q_{i_{1}}(s) \right] \right\}^{1/2} E_{I}(s)^{1/2}.$$

Because

$$s^{2}\operatorname{tr}\left[\left(s^{2}+\omega\right)^{-1}Q_{i}(s)Q_{i}(s)\right] \leq D(s),\tag{7.8}$$

we conclude (i).

From Lemma 6.3 and (6.8), we obtain that

$$||Q_i(s)|| \le \frac{\sqrt{2}}{\nu} ||\hat{\varrho}||_*.$$

Hence, $E_I(s) \leq a^{2j-2}s^2 \operatorname{tr} \left[(s^2 + \omega)^{-1} Q_{i_{2j}}(s) Q_{i_{2j}}(s) \right] \leq a^{2j-2} D(s)$ by (7.8). Therefore, we obtain (7.7).

One observes that

$$\sum_{j=1}^{\infty} \sum_{I \in \mathcal{I}_{2j}^{(e)}} s^2 \left| \operatorname{tr} \left[\left(s^2 + \omega \right)^{-1} Q_I(s) \right] \right| \leq \sum_{\substack{(7.6)\\(7.7)}}^{\infty} \sum_{j=1}^{\infty} \sum_{I \in \mathcal{I}_{2j}^{(e)}} D(s)^{1/2} E_I(s)^{1/2} \leq \sum_{j=1}^{\infty} 2^{2j} a^{j-1} D(s) + \sum_{\substack{(7.7)\\(7.7)}}^{\infty} \frac{4}{1 - 4a} D(s).$$

In the second inequality, we have used the fact that $\#\mathcal{I}_{2j}^{(e)} \leq 2^{2j}$ Accordingly, we get

$$\sum_{j=1}^{\infty} \sum_{I \in \mathcal{I}_{2j}^{(e)}} |\langle Q_I \rangle| \le \frac{4}{1-4a} \frac{1}{2\pi} \int_{\mathbb{R}} \mathrm{d}s \, D(s) \le \frac{4}{1-4a} \frac{e}{\nu} \|\hat{\varrho}\|_*^2$$

by (6.7).

Corollary 7.5. If $\sqrt{2}\|\hat{\varrho}\|_* < \nu$, $1 \leq \sqrt{2}e\nu_0$ and $\sqrt{2}e\|\hat{\varrho}\|_* < 1$, then the r.h.s. of (7.4) converges absolutely for every R > 0. In addition, to exchange the series with the integral, i.e., $\langle \cdots \rangle$ in (7.4) (or (7.3)) can be justified.

8 Proof of Theorem 2.1

For each $I \in \mathcal{I}_{2j}^{(e)}$, #I indicates the cardinality of I. Notice that #I is different from $|I| = i_1 + \cdots + i_{2j}$.

8.1 Analysis of $\langle Q_I \rangle$ with #I = 2

We claim that

$$\langle Q_1 Q_2 \rangle = \langle Q_2 Q_1 \rangle = 0. \tag{8.1}$$

To see this, let $I = \{1, 2\}$ or $\{2, 1\}$. Trivially, |I| = 1 + 2 = 3. By Lemma 7.3, we conclude (8.1).

8.2 Analysis of $\langle Q_I \rangle$ with #I = 4

In this subsection, we will examine the following terms:

$$\sum_{\substack{i_1, \dots, i_4 \in \{1, 2\}\\\{i_1, \dots, i_4\} \neq \{1, 1, 1, 1\}, \{2, 2, 2, 2\}}} \langle Q_{i_1} Q_{i_2} Q_{i_3} Q_{i_4} \rangle = \mathscr{A} + \mathscr{B},$$

where

$$\mathscr{A} = \langle Q_1 Q_1 Q_2 Q_2 \rangle + \langle Q_2 Q_2 Q_1 Q_1 \rangle \tag{8.2}$$

and

$$\mathscr{B} = \langle Q_1 Q_2 Q_1 Q_2 \rangle + \langle Q_2 Q_1 Q_2 Q_1 \rangle + \langle Q_2 Q_1 Q_1 Q_2 \rangle + \langle Q_1 Q_2 Q_2 Q_1 \rangle.$$

$$(8.3)$$

In Appendix B, we will prove the following lemmas.

Lemma 8.1. We have

$$\lim_{R \to \infty} \lim_{\Lambda \to \infty} \lim_{L \to \infty} R^7 \mathscr{A} = \frac{23}{4\pi} \left(\frac{1}{2\pi}\right)^2 \left(\frac{1}{4}\alpha_{\mathrm{E,at}}\right)^2.$$

Lemma 8.2. We have

$$\lim_{R \to \infty} \lim_{\Lambda \to \infty} \lim_{L \to \infty} R^9 \langle Q_1 Q_2 Q_2 Q_1 \rangle = \lim_{R \to \infty} \lim_{\Lambda \to \infty} \lim_{L \to \infty} R^9 \langle Q_2 Q_1 Q_1 Q_2 \rangle = \frac{g}{e^2 \nu^6}$$

where g is a constant independent of e, ν_0 and R. Moreover, $\langle Q_1 Q_2 Q_1 Q_2 \rangle = \langle Q_2 Q_1 Q_2 Q_1 \rangle = 0$. Thus, $\lim_{R \to \infty} \lim_{L \to \infty} \lim_{L \to \infty} R^9 \mathscr{B} = 2g/e^2 \nu^6$.

8.3 Analysis of $\langle Q_I \rangle$ with $\#I \geq 6$

Let $I = \{i_1, \ldots, i_{2j}\} \in \mathcal{I}_{2j}^{(e)}$. We will examine the following two cases, separately.

Case 1: There exists a unique number $\ell \in \{1, 2, ..., 2j - 1\}$ such that $i_{\ell} + i_{\ell+1} = 3$. Case 2: There exist at least two numbers $m, n \in \{1, 2, ..., 2j - 1\}$ such that $i_m + i_{m+1} = i_n + i_{n+1} = 3$.

Example 8.3. For readers' convenience, we provide some examples below:

Case 1:
$$I = \{1, 1, \underbrace{1, 2}^{i_3+i_4=3}, 2, 2, 2, 2\}, \{1, 1, 1, 1, \underbrace{1, 2}^{i_5+i_6=3}, 2, 2\}.$$

Case 2: $I = \{1, 1, \underbrace{1, 2}^{i_3+i_4=3}, 2, 2, 2, 2, \underbrace{2, 1}^{i_8+i_9=3}, 1\}, \{1, 1, 1, 1, 1, \underbrace{1, 2}^{i_6+i_7=3}, \underbrace{2, 1}^{i_8+i_9=3}, 1\}.$

8.3.1 Case 1

In Appendix B, we will prove the following lemma.

Lemma 8.4. Assume that I satisfies the condition in Case 1. If R is sufficiently large, then we have

$$\lim_{\Lambda \to \infty} \lim_{L \to \infty} |\langle Q_I \rangle| \le R^{-9} \alpha_{\mathrm{E,at}}^2 \left(\frac{\|\hat{\varrho}\|_{L^2}^2}{3\nu^2}\right)^{\#I/2-2} C,$$

where C is a positive number independent of e, I, R and ν_0 .

8.3.2 Case 2

The purpose here is to prove Lemma 8.6 below. To this end, we begin with the following lemma.

Lemma 8.5. Let $G = (s^2 + \Omega_0)^{-1}$. For each $j \in \{1, 2\}$, we have

 $\|Q_j G\| \le D(\hat{\varrho}),$

where $D(\hat{\varrho}) = \max\left\{\sqrt{2}e \||k|^{-1}\hat{\varrho}\|_{*}, \frac{\sqrt{2}}{e\nu^{2}}\||k|\hat{\varrho}\|_{*}\right\}.$

Proof. By (A.3) and (A.4), we readily show that

$$\begin{aligned} \left\| \mathbb{T}(r) \left(s^2 + S_0 \right)^{-1} \right\| &\leq \sqrt{2} \left\| |k|^{-1} \hat{\varrho} \right\|_*, \\ \left\| \left(s^2 + e^2 \nu^2 \right)^{-1} \mathbb{T}^*(r) \right\| &\leq \frac{\sqrt{2}}{e^2 \nu^2} \| |k| \hat{\varrho} \|_*. \end{aligned}$$

This concludes the proof of Lemma 8.5.

Lemma 8.6. Let $j \geq 3$. For each $I \in \mathcal{I}_{2j}^{(e)}$ satisfying the condition in Case 2, we have

$$|\langle Q_I \rangle| \le c_L^{2j-4} \langle Q_2 Q_1 Q_1 Q_2 \rangle,$$

where $c_L = \max \left\{ D(\hat{\varrho}), \frac{\sqrt{2}}{\nu} \| \hat{\varrho} \|_* \right\}.$

Proof. By the assumption in the condition Case 2, there exist at least two numbers $m, n \in \{1, 2, \ldots, 2j - 1\}$ such that $i_m + i_{m+1} = i_n + i_{n+1} = 3$. Hence, I can be decomposed as $I = A \cup \{i_m, i_{m+1}\} \cup B \cup \{i_n, i_{n+1}\} \cup C$. Without loss of generality, we may assume that $\{i_m, i_{m+1}\} = \{i_n, i_{n+1}\} = \{1, 2\}$. Thus,

$$\langle Q_I \rangle = \langle Q_A Q_1 Q_2 Q_B Q_1 Q_2 Q_C \rangle.$$

Let $Q_I(s) = Q_{i_1}(s)Q_{i_2}(s)\cdots Q_{i_{2j}}(s)$. By the Schwarz inequality, we have

$$\left| \operatorname{tr} \left[G^{1/2} Q_I(s) G^{1/2} \right] \right| \le \Phi_1^{1/2} \Phi_2^{1/2},$$
(8.4)

where

$$\Phi_1 = \operatorname{tr} \left[G^{1/2} Q_A(s) Q_1(s) Q_2(s) Q_2(s) Q_1(s) Q_A^*(s) G^{1/2} \right],$$

$$\Phi_2 = \operatorname{tr} \left[G^{1/2} Q_C^*(s) Q_2(s) Q_1(s) Q_B^*(s) Q_B(s) Q_1(s) Q_2(s) Q_C(s) G^{1/2} \right].$$

First, we estimate Φ_1 . By the cyclic property of the trace, we have

$$\Phi_{1} = \operatorname{tr} \left[Q_{2}(s)Q_{1}(s)Q_{A}^{*}(s)G^{1/2}G^{1/2}Q_{A}(s)Q_{1}(s)Q_{2}(s) \right]$$

=
$$\operatorname{tr} \left[Q_{2}(s)Q_{1}(s)G^{1/2}G^{-1/2}Q_{A}^{*}(s)G^{1/2}G^{1/2}Q_{A}(s)G^{-1/2}G^{1/2}Q_{1}(s)Q_{2}(s) \right].$$
(8.5)

Because

$$G^{1/2}Q_A(s)G^{-1/2} = (GQ_{a_1})(GQ_{a_2})\cdots(GQ_{a_{\#A}})$$

where $A = \{a_1, a_2, ..., a_{\#A}\}$, we have, by Lemma 8.5,

$$\left\| G^{1/2} Q_A(s) G^{-1/2} \right\| \le D(\hat{\varrho})^{\# A}.$$

Thus, by (8.5) and the cyclic property of the trace,

$$\Phi_1 \le D(\hat{\varrho})^{2\#A} \operatorname{tr} \left[G^{1/2} Q_1(s) Q_2(s) Q_2(s) Q_1(s) G^{1/2} \right].$$
(8.6)

As for Φ_2 , we have

$$\Phi_2 \le \|Q_B(s)\|^2 \operatorname{tr} \left[G^{1/2} Q_C^*(s) Q_2(s) Q_1(s) Q_1(s) Q_2(s) Q_C(s) G^{1/2} \right].$$

By an argument similar to the one in the proof of (8.6), one obtains that

$$\operatorname{tr} \left[G^{1/2} Q_C^*(s) Q_2(s) Q_1(s) Q_1(s) Q_2(s) Q_C(s) G^{1/2} \right] \\ \leq D(\hat{\varrho})^{2 \# C} \operatorname{tr} \left[G^{1/2} Q_2(s) Q_1(s) Q_1(s) Q_2(s) G^{1/2} \right].$$
(8.7)

By using the fact $||Q_B(s)|| \leq \left(\frac{\sqrt{2}}{\nu} ||\hat{\varrho}||_*\right)^{2\#B}$ and (8.7), we have

$$\Phi_2 \le \left(\frac{\sqrt{2}}{\nu} \|\hat{\varrho}\|_*\right)^{2\#B} D(\hat{\varrho})^{2\#C} \operatorname{tr} \left[G^{1/2} Q_2(s) Q_1(s) Q_1(s) Q_2(s) G^{1/2}\right].$$
(8.8)

Combining (8.4), (8.6) and (8.8), we arrive at

$$\begin{split} |\langle Q_I \rangle| &\leq \left(\frac{\sqrt{2}}{\nu} \|\hat{\varrho}\|_*\right)^{\#B} D(\hat{\varrho})^{\#A+\#C} \langle Q_1 Q_2 Q_2 Q_1 \rangle^{1/2} \langle Q_2 Q_1 Q_1 Q_2 \rangle^{1/2} \\ &\leq c_L^{2j-4} \langle Q_1 Q_2 Q_2 Q_1 \rangle^{1/2} \langle Q_2 Q_1 Q_1 Q_2 \rangle^{1/2}. \end{split}$$

Because $\langle Q_1 Q_2 Q_2 Q_1 \rangle = \langle Q_2 Q_1 Q_1 Q_2 \rangle$, we obtain the desired result.

8.4 Completion of the proof of Theorem 2.1

First, remark that $\lim_{\Lambda \to \infty} \lim_{L \to \infty} E_{L,\Lambda} = E$ and $\lim_{\Lambda \to \infty} \lim_{L \to \infty} E_{L,\Lambda}(R) = E(R)$ by Proposition 5.1. We divide $\mathcal{I}_{2j}^{(e)}$ as $\mathcal{I}_{2j}^{(e)} = \mathcal{I}_{2j,1}^{(e)} \cup \mathcal{I}_{2j,2}^{(e)}$, where

 $\mathcal{I}_{2j,\alpha}^{(e)} = \big\{ I \in \mathcal{I}_{2j}^{(e)} \, | \, I \text{ satisfies the condition in Case } \alpha \big\}, \qquad \alpha = 1, 2.$

Note that $\#\mathcal{I}_{2j,1}^{(e)} = 2j - 1$ and $\#\mathcal{I}_{2j,2}^{(e)} \le 2^{2j}$. By (7.5) and (8.1), one obtains that

$$2E_{L,\Lambda} - E_{L,\Lambda}(R) = \mathscr{A} + \mathscr{B} + \sum_{j\geq 3} \sum_{I\in\mathcal{I}_{2j,1}^{(e)}} \langle Q_I \rangle + \sum_{j\geq 3} \sum_{I\in\mathcal{I}_{2j,2}^{(e)}} \langle Q_I \rangle,$$

where \mathscr{A} and \mathscr{B} are defined by (8.2) and (8.3), respectively. Therefore,

$$\left|R^{7}\left\{2E_{L,\Lambda}-E_{L,\Lambda}(R)-\mathscr{A}\right\}\right| \leq R^{7}\mathscr{B}+\sum_{j\geq3}\sum_{I\in\mathcal{I}_{2j,1}^{(e)}}R^{7}|\langle Q_{I}\rangle|+\sum_{j\geq3}\sum_{I\in\mathcal{I}_{2j,2}^{(e)}}R^{7}|\langle Q_{I}\rangle|.$$
(8.9)

We will estimate the three terms in the right-hand side of (8.9). By Lemma 8.2, we can easily control the first term. As for the second term, by Lemma 8.4, we have

$$\lim_{\Lambda \to \infty} \lim_{L \to \infty} \sum_{j \ge 3} \sum_{I \in \mathcal{I}_{2j,1}^{(e)}} R^7 |\langle Q_I \rangle| \leq \lim_{\text{Lemma 8.4}} R^{-2} \sum_{j \ge 3} \left(\# \mathcal{I}_{2j,1}^{(e)} \right) C \alpha_{\text{E,at}}^2 \left(\frac{\|\hat{\varrho}\|_{L^2}^2}{3\nu^2} \right)^{j-2} = R^{-2} C \alpha_{\text{E,at}}^2 \sum_{j \ge 3} (2j-1) \left(\frac{\|\hat{\varrho}\|_{L^2}^2}{3\nu^2} \right)^{j-2}.$$
(8.10)

Note that because $\|\hat{\varrho}\|_{L^2}^2/3\nu^2 < 1$, the right-hand side of (8.10) converges. On the other hand, using Lemma 8.6, one obtains that

$$\sum_{j\geq3} \sum_{I\in\mathcal{I}_{2j,2}^{(e)}} R^{7} |\langle Q_{I}\rangle| \leq_{\text{Lemma 8.6}} \sum_{j\geq3} \left(\#\mathcal{I}_{2j,2}^{(e)} \right) c_{L}^{2j-4} R^{7} \langle Q_{2}Q_{1}Q_{1}Q_{2}\rangle \\ \leq \sum_{j\geq3} 2^{2j} c_{L}^{2j-4} R^{7} \langle Q_{2}Q_{1}Q_{1}Q_{2}\rangle.$$

$$(8.11)$$

Note that because $\lim_{L\to\infty} c_L = c_{\infty} < 1/2$, the right-hand side of (8.11) converges, provided that L is sufficiently large.

Combining (8.9), (8.10) and (8.11), and using Lemma 8.2, we finally arrive at

$$\begin{split} \lim_{R \to \infty} \left| R^7 \left\{ 2E - E(R) - \frac{23}{4\pi} \left(\frac{1}{2\pi} \right)^2 \left(\frac{1}{4} \alpha_{\mathrm{E,at}} \right)^2 \right\} \right| \\ &\leq \left\{ 2 + c_{\infty}^{-4} \sum_{j \ge 3} (2c_{\infty})^{2j} \right\} \lim_{R \to \infty} \lim_{\Lambda \to \infty} \lim_{L \to \infty} R^7 \langle Q_1 Q_2 Q_2 Q_1 \rangle \\ &+ \lim_{R \to \infty} R^{-2} C \alpha_{\mathrm{E,at}}^2 \sum_{j \ge 3} (2j-1) \left(\frac{\|\hat{\varrho}\|_{L^2}^2}{3\nu^2} \right)^{2j-2} = 0. \end{split}$$

This concludes the proof of Theorem 2.1.

9 Discussions

9.1 Indistinguishability of the electrons

The original Hamiltonian H_{2e} has the indistinguishability of the electrons, i.e., the Hamiltonian is unchanged under the exchange of $x_1 \leftrightarrow x_2 + r$. In contrast to this, the approximated Hamiltonian H_{D2e} breaks the indistinguishability. Nevertheless, the Hamiltonian H_{D2e} does explain the Casimi–Polder potential as we show in Theorem 2.1. The distinguishability comes from the assumptions (C.1) and (C.2). However, to justify the assumptions is still open.

One way to avoid the unjustified derivation of H_{D2e} is to directly start with the Hamiltonian H given by (4.3) without the last term, which can for instance be directly taken from [24, equation (13.127)] and then extended to the two-particle case. Alternatively and equivalently, the many-particle case is presented, e.g., in [17, Section 4]. If we start from this form, the necessary assumptions are stated as follows:

- We assume distinguishability of the two electrons by localizing electron 1 at 0, such that electron 1 experiences the field $\hat{E}(0)$, while electron 2 is localized at r and hence experiences the field $\hat{E}(r)$.
- We discard all self-interaction terms and approximate the atomic Coulomb potential by a harmonic potential.

In this manner, we can construct a minimal QED model which describes the Casimir–Polder potential. Note that, since the particle 1 and 2 only communicate via the photon field, and due to distinguishability, the actual choice of coordinate systems is insubstantial such that we can choose for particle 2 a coordinate system that is centered at r.

9.2 Cancellation mechanism of the van der Waals–London force

As we performed in [20], the attractive R^{-7} decay (the retarded van der Waals potential) appears due to the exact cancellation of the terms with R^{-6} decay (the van der Waals–London potential) originating from V_R by the contribution from the quantized radiation field. Note that the conditions (A.1)–(A.3) are assumed in [20] as well, but (C.1) and (C.2) are not. As we saw in the present paper, this kind of the cancellation mechanism cannot be reproduced under the conditions (C.1), (C.2) and (A.1)–(A.4). In this sense, our assumptions, especially (C.1) and (C.2) would be unphysical.

In many literatures, the retardation on the van der Waals potential is examined under the condition (C.1) alone. In these studies, the cancellation of the terms with R^{-6} decay is presupposed and only the 4-th order perturbation theory is performed without estimating higher order terms.⁶ As far as we know, to examine the exact cancellation mechanism under only the condition (C.1) is still unsolved. This problem could be a key to achieving mathematically complete understanding of the retarded van der Waals potential.

A Useful formulas

In this appendix, we give a list of useful formulas. Let $\mathbb{T}_s(x) = \mathbb{T}(x)(s^2 + S_0)^{-1/2}$. First, we give some formulas for Q_i :

$$Q_{1}(s) = e \left(s^{2} + e^{2}\nu^{2}\right)^{-1/2} \begin{pmatrix} 0 & 0 & \mathbb{T}_{s}(0) \\ 0 & 0 & 0 \\ \mathbb{T}_{s}^{*}(0) & 0 & 0 \end{pmatrix},$$

$$Q_{2}(s) = e \left(s^{2} + e^{2}\nu^{2}\right)^{-1/2} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & \mathbb{T}_{s}(r) \\ 0 & \mathbb{T}_{s}^{*}(r) & 0 \end{pmatrix},$$

$$Q_{1}(s)^{2} = e^{2} \left(s^{2} + e^{2}\nu^{2}\right)^{-1} \begin{pmatrix} \mathbb{T}_{s}(0)\mathbb{T}_{s}^{*}(0) & 0 & 0 \\ 0 & 0 & \mathbb{T}_{s}^{*}(0)\mathbb{T}_{s}(0) \end{pmatrix},$$
(A.1)

$$Q_{2}(s)^{2} = e^{2} \left(s^{2} + e^{2} \nu^{2}\right)^{-1} \begin{pmatrix} 0 & 0 & 0 \\ 0 & \mathbb{T}_{s}(r) \mathbb{T}_{s}^{*}(r) & 0 \\ 0 & 0 & \mathbb{T}_{s}^{*}(r) \mathbb{T}_{s}(r) \end{pmatrix},$$
(A.2)
$$Q_{1}(s)Q_{2}(s) = e^{2} \left(s^{2} + e^{2} \nu^{2}\right)^{-1} \begin{pmatrix} 0 & \mathbb{T}_{s}(0) \mathbb{T}_{s}^{*}(r) & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix},$$
$$Q_{2}(s)Q_{1}(s) = e^{2} \left(s^{2} + e^{2} \nu^{2}\right)^{-1} \begin{pmatrix} 0 & 0 & 0 \\ \mathbb{T}_{s}(r) \mathbb{T}_{s}^{*}(0) & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}.$$

Let $\mathbb{M}(r, r')$ be a linear operator on $\ell^2(M \times \{1, \ldots, 4\})$ defined by

$$(\mathbb{M}(r,r')\mathbf{f})(k,\lambda) = \sum_{\lambda'=1}^{4} \sum_{k'\in M} M_{k,\lambda;k',\lambda'}(r,r')f(k',\lambda'), \qquad \mathbf{f} \in \ell^2(M \times \{1,\dots,4\}),$$
$$M_{k,\lambda;k',\lambda'}(r,r') = \langle \varepsilon(k,\lambda) | \varepsilon(k',\lambda') \rangle_3 F_r(k,\lambda) F_{r'}(k',\lambda').$$

The following formulas are readily checked:

$$\mathbb{T}_{s}^{*}(r)\mathbb{T}_{s}(r') = \left(s^{2} + S_{0}\right)^{-1/2}\mathbb{M}(r, r')\left(s^{2} + S_{0}\right)^{-1/2},\tag{A.3}$$

⁶A kind of weak cancellation mechanics is discussed in [12] by imposing the infrared cutoff.

$$\mathbb{T}_{s}(r)\mathbb{T}_{s}^{*}(r') = \sum_{\lambda=1}^{4} \sum_{k \in M} |\varepsilon(k,\lambda)\rangle \langle \varepsilon(k,\lambda)| (s^{2} + k^{2})^{-1} F_{r}(k,\lambda) F_{r'}(k,\lambda).$$
(A.4)

Note that $\mathbb{T}_s^*(r)\mathbb{T}_s(r')$ is a map from $\ell^2(M \times \{1, \ldots, 4\})$ to $\ell^2(M \times \{1, \ldots, 4\})$, while $\mathbb{T}_s(r)\mathbb{T}_s^*(r')$ is a map from \mathbb{C}^3 to \mathbb{C}^3 .

B Numerical computations

B.1 Proof of Lemma 8.1

We will extend the methods in [20, 21]. By (A.1) and (A.2), we have

$$Q_1(s)^2 Q_2(s)^2 = e^4 \left(s^2 + e^2 \nu^2\right)^{-2} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & \mathbb{T}_s^*(0) \mathbb{T}_s(0) \mathbb{T}_s^*(r) \mathbb{T}_s(r) \end{pmatrix},$$

which implies that

$$\langle Q_1 Q_1 Q_2 Q_2 \rangle = \frac{e^4}{2\pi} \int_{\mathbb{R}} \mathrm{d}s \, \frac{s^2}{\left(s^2 + e^2 \nu^2\right)^2} \, \mathrm{tr} \left[\mathbb{T}_s^*(0) \mathbb{T}_s(0) \mathbb{T}_s^*(r) \mathbb{T}_s(r)\right]$$

By (A.3) and the fact $\sum_{\lambda_1,\lambda_2=1,2} (\langle \varepsilon(k_1,\lambda_1) | \varepsilon(k_2,\lambda_2) \rangle_3)^2 = 1 + (\hat{k}_1 \cdot \hat{k}_2)^2$ with $\hat{k} = k/|k|$, we have

$$\operatorname{tr} \left[\mathbb{T}_{s}^{*}(0)\mathbb{T}_{s}(0)\mathbb{T}_{s}^{*}(r)\mathbb{T}_{s}(r) \right] = \operatorname{tr} \left[\left(s^{2} + \Omega_{0} \right)^{-2} \mathbb{M}(0,0) \left(s^{2} + \Omega_{0} \right)^{-1} \mathbb{M}(r,r) \right]$$

$$= \sum_{k_{1},\lambda_{1}} \sum_{k_{2},\lambda_{2}} \left(s^{2} + k_{1}^{2} \right)^{-2} \left(s^{2} + k_{2}^{2} \right)^{-1} \left(\langle \varepsilon(k_{1},\lambda_{1}) | \varepsilon(k_{2},\lambda_{2}) \rangle_{3} \right)^{2}$$

$$\times k_{1}^{2} k_{2}^{2} \hat{\varrho}(k_{1})^{2} \hat{\varrho}(k_{2})^{2} \cos(k_{1} \cdot r) \cos(k_{2} \cdot r)$$

$$= \sum_{k_{1}} \sum_{k_{2}} \left(s^{2} + k_{1}^{2} \right)^{-2} \left(s^{2} + k_{2}^{2} \right)^{-1} \left(1 + \left(\hat{k}_{1} \cdot \hat{k}_{2} \right)^{2} \right) k_{1}^{2} k_{2}^{2} \hat{\varrho}(k_{1})^{2} \hat{\varrho}(k_{2})^{2} \cos(k_{1} \cdot r) \cos(k_{2} \cdot r).$$

$$(B.1)$$

By using the assumption (A.3), we have

the r.h.s. of (B.1)
=
$$\sum_{k_1} \sum_{k_2} (s^2 + k_1^2)^{-2} (s^2 + k_2^2)^{-1} (1 + (\hat{k}_1 \cdot \hat{k}_2)^2) k_1^2 k_2^2 \hat{\varrho}(k_1)^2 \hat{\varrho}(k_2)^2 \cos((k_1 + k_2) \cdot r).$$

Hence, we arrive at

$$\begin{aligned} \langle Q_1 Q_1 Q_2 Q_2 \rangle &= \frac{e^4}{2} \left(\frac{2\pi}{L} \right)^6 \sum_{k_1, k_2 \in M} |k_1|^2 |k_2|^2 \left(1 + \left(\hat{k}_1 \cdot \hat{k}_2 \right)^2 \right) \hat{\varrho}^2(k_1) \hat{\varrho}^2(k_2) \cos\{ (k_1 + k_2) \cdot r \} \\ &\times I_{2,2,1} \left(e^2 \nu^2; |k_1|^2; |k_2|^2 \right) \end{aligned}$$

where

$$I_{n_a,n_b,n_c}(a;b;c) = \frac{1}{\pi} \int_{-\infty}^{\infty} \mathrm{d}s \, \frac{s^2}{\left(s^2 + a\right)^{n_a} \left(s^2 + b\right)^{n_b} \left(s^2 + c\right)^{n_c}}.$$
(B.2)

Thus, we obtain that

$$\lim_{\Lambda \to \infty} \lim_{L \to \infty} \langle Q_1 Q_1 Q_2 Q_2 \rangle = \frac{e^4}{2} \int_{\mathbb{R}^3 \times \mathbb{R}^3} \mathrm{d}k_1 \mathrm{d}k_2 |k_1|^2 |k_2|^2 \left(1 + \left(\hat{k}_1 \cdot \hat{k}_2\right)^2\right) \hat{\varrho}^2(k_1) \hat{\varrho}^2(k_2)$$

× cos{
$$(k_1 + k_2) \cdot r$$
} $I_{2,2,1}(e^2\nu^2; |k_1|^2; |k_2|^2).$ (B.3)

By scalings $Rk_1 \rightsquigarrow k_1$ and $Rk_2 \rightsquigarrow k_2$, we have

the r.h.s. of (B.3) =
$$R^{-10} \frac{e^4}{2} \int dk_1 dk_2 |k_1|^2 |k_2|^2 (1 + (\hat{k}_1 \cdot \hat{k}_2)^2) \hat{\varrho}^2 (k_1/R) \hat{\varrho}^2 (k_2/R) e^{i(k_1+k_2)\cdot \hat{n}} \times I_{2,2,1} (e^2 \nu^2; |k_1|^2/R^2; |k_2|^2/R^2),$$
 (B.4)

where $\hat{n} = r/R = (0, 0, 1)$. Let us switch to spherical coordinates (r, φ, θ) by

$$\hat{k} = (Y \cos \varphi, Y \sin \varphi, X), \qquad X = \cos \theta, \qquad Y = \sin \theta.$$

Clearly $X^2 + Y^2 = 1$. Then we have

$$\hat{k}_1 \cdot \hat{k}_2 = \cos(\varphi_1 - \varphi_2)Y_1Y_2 + X_1X_2$$

and hence, by taking the symmetry between r_1 and r_2 variables into consideration, we obtain

the r.h.s. of (B.4) =
$$R^{-10} \frac{e^4}{2} \int_0^\infty \mathrm{d}r_1 \int_0^\infty \mathrm{d}r_2 \int_{-1}^1 \mathrm{d}X_1 \int_{-1}^1 \mathrm{d}X_2 \,\mathfrak{S}(X_1, X_2)$$

 $\times r_1^4 r_2^4 \mathrm{e}^{\mathrm{i}r_1 X_1} \mathrm{e}^{\mathrm{i}r_2 X_2} \mathbf{I} \left(e^2 \nu^2; r_1^2/R^2; r_2^2/R^2 \right) \hat{\varrho}_{\mathrm{rad}}^2(r_1/R) \hat{\varrho}_{\mathrm{rad}}^2(r_2/R),$

where

$$\mathfrak{S}(X_1, X_2) = \int_0^{2\pi} \mathrm{d}\varphi_1 \int_0^{2\pi} \mathrm{d}\varphi_2 \left\{ 1 + \left(\cos(\varphi_1 - \varphi_2) Y_1 Y_2 + X_1 X_2 \right)^2 \right\} \\ = 6\pi^2 - 2\pi^2 \left(X_1^2 + X_2^2 \right) + 6\pi^2 X_1^2 X_2^2.$$
(B.5)

 $\quad \text{and} \quad$

$$\mathbf{I}(a;b;c) = \frac{1}{2} \left\{ I_{2,2,1}(a;b;c) + I_{2,1,2}(a;b;c) \right\}$$

By (C.2) and (C.3), we decompose I(a; b; c) as

$$\mathbf{I}(a;b;c) = \mathbf{I}_{\rm re}(a;b;c) + \mathbf{I}_{\rm ir}(a;b;c),$$

where

$$\begin{aligned} \mathbf{I}_{\rm re}(a;b;c) &= \frac{I_{1,1,1}(a;b;c)}{8\sqrt{abc}} \left(\frac{1}{A} + \frac{1}{C}\right),\\ \mathbf{I}_{\rm ir}(a;b;c) &= \frac{I_{1,1,1}(a;b;c)}{8\sqrt{ab}} \frac{1}{A} \left(\frac{2}{A} + \frac{1}{C}\right) + \frac{I_{1,1,1}(a;b;c)}{8\sqrt{ac}} \frac{1}{C} \left(\frac{2}{C} + \frac{1}{A}\right) \end{aligned}$$

with $A = \sqrt{a} + \sqrt{b}$, $B = \sqrt{b} + \sqrt{c}$ and $C = \sqrt{c} + \sqrt{a}$. First, we compute the contribution from the term \mathbf{I}_{re} . By the formula

$$\mathbf{I}_{\rm re} \left(e^2 \nu^2; r_1^2 / R^2; r_2^2 / R^2 \right) = \frac{R^3}{8e\nu r_1 r_2 (e\nu + r_1/R) (e\nu + r_2/R)} \left(\frac{1}{e\nu + r_1/R} + \frac{1}{e\nu + r_2/R} \right) \int_0^\infty \mathrm{d}t \, \mathrm{e}^{-t(r_1 + r_2)},$$

the contribution can be expressed as

$$\frac{e^3}{16\nu}R^{-7}\int_0^\infty dt \int_0^\infty dr_1 \int_0^\infty dr_2 \left(\frac{1}{e\nu + r_1/R} + \frac{1}{e\nu + r_2/R}\right)$$

$$\times \left\{ \frac{3}{2} [1]_{r_1} [1]_{r_2} - [1]_{r_1} [X^2]_{r_2} + \frac{3}{2} [X^2]_{r_1} [X^2]_{r_2} \right\},\tag{B.6}$$

where

$$[A(X)]_r(t) = (2\pi) \frac{r^3 \mathrm{e}^{-tr} \hat{\varrho}_{\mathrm{rad}}^2(r/R)}{e\nu + r/R} \int_{-1}^1 \mathrm{d}X \, \mathrm{e}^{\mathrm{i}rX} A(X).$$

For readers' convenience, we will explain how to compute the integral (B.6). Let $\varphi(x)$ be the Fourier transformation of $\hat{\varrho}_{\rm rad}^2(r)$: $\varphi(x) = (2\pi)^{-1/2} \int_{\mathbb{R}} e^{-irx} \hat{\varrho}_{\rm rad}^2(r) dr$. Here, we extend $\hat{\varrho}_{\rm rad}^2$ to a function on \mathbb{R} by $\hat{\varrho}_{\rm rad}^2(-r) := \hat{\varrho}_{\rm rad}^2(r)$ for r > 0. Note that $\varphi(x)$ decays rapidly by the assumption (A.3). By the convolution theorem in the Fourier analysis, we have

$$\int_{0}^{\infty} \mathrm{d}r \, [1]_{r} = \int_{0}^{\infty} \mathrm{d}s \int_{\mathbb{R}} \mathrm{d}x \, (1+x/R) \mathrm{e}^{-se\nu} \frac{12(t+s/R)^{2} - 4(1+x/R)^{2}}{\left\{(t+s/R)^{2} + (1+x/R)^{2}\right\}^{3}} \varphi(x) \tag{B.7}$$

and

$$\int_0^\infty \mathrm{d}r \, \frac{[1]_r}{e\nu + r/R} = \int_0^\infty \mathrm{d}s \int_{\mathbb{R}} \mathrm{d}x \, (1 + x/R) s \mathrm{e}^{-se\nu} \frac{12(t + s/R)^2 - 4(1 + x/R)^2}{\left\{(t + s/R)^2 + (1 + x/R)^2\right\}^3} \varphi(x) dx$$

[Here, we explain how we derive (B.7). First, we observe that

$$\int_{0}^{\infty} dr [1]_{r} = 4\pi \int_{0}^{\infty} dr \, \frac{r^{2} e^{-tr} \hat{\varrho}_{rad}^{2}(r/R)}{e\nu + r/R} \sin r$$
$$= 4\pi \int_{0}^{\infty} ds \, e^{-se\nu} \operatorname{Im} \int_{\mathbb{R}} dr \underbrace{1_{+}(r)r^{2} e^{-(t+s/R)r}}_{=:f(r)} \hat{\varrho}_{rad}^{2}(r/R) e^{ir},$$

where $1_+(r) = 1$ if r > 0, $1_+(r) = 0$ if $r \le 0$. By the convolution theorem $((2\pi)^{1/2} (\hat{g}\hat{h})^{\vee} = g * h)$, we have

$$(2\pi)^{1/2}\operatorname{Im}\int_{\mathbb{R}} \mathrm{d}r f(r)\hat{\varrho}_{\mathrm{rad}}^{2}(r/R)\mathrm{e}^{\mathrm{i}rx} = \left(\varphi_{R} * \operatorname{Im}\check{f}\right)(x),$$

where $\varphi_R(x) = R\varphi(Rx)$ and $(g * h)(x) = \int_{\mathbb{R}} g(y)h(x - y)dy$. Because

Im
$$\check{f}(x) = (2\pi)^{-1/2} \frac{6x(t+s/R)^2 - 2x^3}{\left\{(t+s/R)^2 + x^2\right\}^3},$$

we get (B.7).] Hence, by the dominated convergence theorem, we obtain

$$\lim_{R \to \infty} \frac{e^3}{16\nu} \int_0^\infty dt \int_0^\infty dr_1 \int_0^\infty dr_2 \left(\frac{1}{e\nu + r_1/R} + \frac{1}{e\nu + r_2/R}\right) [1]_{r_1} [1]_{r_2}$$
$$= \frac{\pi}{8\nu^4} \hat{\varrho}^4(0) \int_0^\infty dt \, \frac{\left(-4 + 12t^2\right)^2}{\left(t^2 + 1\right)^6}.$$

Similarly, we obtain that

$$\lim_{R \to \infty} \frac{e^3}{16\nu} \int_0^\infty dt \int_0^\infty dr_1 \int_0^\infty dr_2 \left(\frac{1}{e\nu + r_1/R} + \frac{1}{e\nu + r_2/R}\right) [1]_{r_1} [X^2]_{r_2}$$
$$= \frac{\pi}{8\nu^4} \hat{\varrho}^4(0) \int_0^\infty dt \, \frac{\left(-4 + 12t^2\right)\left(-12 + 4t^2\right)}{\left(t^2 + 1\right)^6}$$

and

$$\lim_{R \to \infty} \frac{e^3}{16\nu} \int_0^\infty dt \int_0^\infty dr_1 \int_0^\infty dr_2 \left(\frac{1}{e\nu + r_1/R} + \frac{1}{e\nu + r_2/R} \right) \left[X^2 \right]_{r_1} \left[X^2 \right]_{r_2}$$
$$= \frac{\pi}{8\nu^4} \hat{\varrho}^4(0) \int_0^\infty dt \, \frac{\left(-12 + 4t^2 \right)^2}{\left(t^2 + 1 \right)^6}.$$

Summarizing the above results, we arrive at

$$\frac{e^4}{2} \int_0^\infty \mathrm{d}r_1 \int_0^\infty \mathrm{d}r_2 \int_{-1}^1 \mathrm{d}X_1 \int_{-1}^1 \mathrm{d}X_2 \,\mathfrak{S}(X_1, X_2) r_1^4 r_2^4 \mathrm{e}^{\mathrm{i}r_1 X_1} \mathrm{e}^{\mathrm{i}r_2 X_2} \\ \times \mathbf{I}_{\mathrm{re}} \big(e^2 \nu^2; r_1^2 / R^2; r_2^2 / R^2 \big) \hat{\varrho}_{\mathrm{rad}}^2 (r_1 / R) \hat{\varrho}_{\mathrm{rad}}^2 (r_2 / R) = \frac{23\pi^3}{2\nu^4} \hat{\varrho}^4(0) R^{-7} + o \big(R^{-7} \big),$$

where we used the following formula in [20]:

$$\int_0^\infty dt \left\{ \frac{3}{2} A(t)^2 - A(t)B(t) + \frac{3}{2}B(t)^2 \right\} = 92\pi^3$$

with

$$A(t) = \frac{-4 + 12t^2}{(1+t^2)^3}$$
 and $B(t) = \frac{4(-3+t^2)}{(1+t^2)^3}.$

As for the contribution from I_{ir} , we have, by an argument similar to that of the computation concerning with I_{re} ,

$$\frac{e^4}{2} \int_0^\infty \mathrm{d}r_1 \int_0^\infty \mathrm{d}r_2 \int_{-1}^1 \mathrm{d}X_1 \int_{-1}^1 \mathrm{d}X_2 \,\mathfrak{S}(X_1, X_2) \, r_1^4 r_2^4 \mathrm{e}^{\mathrm{i}r_1 X_1} \mathrm{e}^{\mathrm{i}r_2 X_2} \\ \times \,\mathbf{I}_{\mathrm{ir}} \big(e^2 \nu^2; r_1^2/R^2; r_2^2/R^2 \big) \hat{\varrho}_{\mathrm{rad}}^2(r_1/R) \hat{\varrho}_{\mathrm{rad}}^2(r_2/R) = \mathrm{const} \cdot R^{-9} + o\big(R^{-9}\big).$$

To summarize, we obtain that

$$\lim_{R \to \infty} \lim_{\Lambda \to \infty} \lim_{L \to \infty} R^7 \langle Q_1 Q_1 Q_2 Q_2 \rangle = \frac{23}{8\pi} \left(\frac{1}{2\pi} \right)^2 \left(\frac{1}{4} \alpha_{\mathrm{E,at}} \right)^2.$$

Similarly, we get

$$\lim_{R \to \infty} \lim_{\Lambda \to \infty} \lim_{L \to \infty} R^7 \langle Q_2 Q_2 Q_1 Q_1 \rangle = \frac{23}{8\pi} \left(\frac{1}{2\pi} \right)^2 \left(\frac{1}{4} \alpha_{\mathrm{E,at}} \right)^2.$$

This concludes the proof of Lemma 8.1.

B.2 Proof of Lemma 8.2

We readily see that $\langle Q_1 Q_2 Q_1 Q_2 \rangle = \langle Q_2 Q_1 Q_2 Q_1 \rangle = 0$ by the formulas in Appendix A. In what follows, we evaluate $\langle Q_1 Q_2 Q_2 Q_1 \rangle$ and $\langle Q_2 Q_1 Q_1 Q_2 \rangle$. Because the argument here is almost pallarel to the proof of Lemma 8.1, we provide a sketch only. As before, we have

$$\lim_{\Lambda \to \infty} \lim_{L \to \infty} \langle Q_2 Q_1 Q_1 Q_2 \rangle = \frac{e^4}{2\pi} \int dk_1 dk_2 \left(1 + \left(\hat{k}_1 \cdot \hat{k}_2 \right)^2 \right) |k_1|^2 |k_2|^2 \hat{\varrho}^2(k_1) \hat{\varrho}^2(k_2) \\ \times \cos(k_1 \cdot r) \cos(k_2 \cdot r) I_{3,1,1} \left(e^2 \nu^2; |k_1|^2; |k_2|^2 \right) \\ = R^{-10} \frac{e^4}{2} \int_0^\infty dr_1 \int_0^\infty dr_2 \int_{-1}^1 dX_1 \int_{-1}^1 dX_2 \,\mathfrak{S}(X_1, X_2) r_1^4 r_2^4 \hat{\varrho}_{\rm rad}^2(r_1/R) \hat{\varrho}_{\rm rad}^2(r_2/R)$$

$$\times \cos(r_1 X_1) \cos(r_2 X_2) I_{3,1,1} \left(e^2 \nu^2; r_1^2 / R^2; r_2^2 / R^2 \right).$$
(B.8)

Remark the following formula:

$$I_{3,1,1}(e^2\nu^2; r_1^2/R^2; r_2^2/R^2) = (e\nu)^{-6} \frac{R}{r_1 + r_2} + o(R),$$

which follows from (C.4). Inserting this into (B.8), we formally obtain that

$$\lim_{\Lambda \to \infty} \lim_{L \to \infty} \langle Q_2 Q_1 Q_1 Q_2 \rangle = \frac{g}{e^2 \nu^6} R^{-9} + o(R^{-9})$$

To justify this rough argument, we carefully have to treat the oscillatory integral as we did in the proof of Lemma 8.1. Similarly, we see that $\lim_{\Lambda \to \infty} \lim_{L \to \infty} \langle Q_1 Q_2 Q_2 Q_1 \rangle = \frac{g}{e^2 \nu^6} R^{-9} + o(R^{-9}).$

B.3 Proof of Lemma 8.4

In this case, there exist two numbers $m, n \in \mathbb{N}$ with $m + n \geq 3$ such that $\langle Q_I \rangle = \langle Q_1^{2m} Q_2^{2n} \rangle$ or $\langle Q_I \rangle = \langle Q_2^{2m} Q_1^{2n} \rangle$. We will study the case where $\langle Q_I \rangle = \langle Q_1^{2m} Q_2^{2n} \rangle$ only. By using the formulas in Appendix A, one obtains that

$$\lim_{\Lambda \to \infty} \lim_{L \to \infty} \left\langle Q_1^{2m} Q_2^{2n} \right\rangle$$

$$= e^{2(m+n)} \sum_{\lambda_1, \dots, \lambda_{m+n} = 1, 2} \int \mathrm{d}k_1 \cdots \mathrm{d}k_{m+n} I_{m+n, 2, 1^{m+n-1}} \left(e^2 \nu^2; |k_1|^2; \dots; |k_{m+n}|^2 \right)$$

$$\times \left[\prod_{j=1}^{m+n} |k_j|^2 \hat{\varrho}^2(k_j) \right] \left\langle \varepsilon_1 | \varepsilon_2 \right\rangle \left\langle \varepsilon_2 | \varepsilon_3 \right\rangle \cdots \left\langle \varepsilon_{m+n} | \varepsilon_1 \right\rangle \cos(k_m \cdot r) \cos(k_{m+1} \cdot r), \quad (B.9)$$

where $\varepsilon_j = \varepsilon(k_j, \lambda_j)$ and $I_{m+n,2,1^{m+n-1}}(a_0; \ldots; a_{m+n}) = I_{m+n,2,\underbrace{1,\ldots,1}_{m+n-1}}(a_0; \ldots; a_{m+n})$ with

$$I_{n_0,n_1,\dots,n_k}(a_0;a_1;\dots;a_k) = \frac{1}{\pi} \int_{-\infty}^{\infty} \mathrm{d}s \, \frac{s^2}{\prod_{j=0}^k (s^2 + a_j)^{n_j}}$$

By scalings $Rk_m \rightsquigarrow k_m$ and $Rk_{m+1} \rightsquigarrow k_{m+1}$, we get

the r.h.s. of (B.9) =
$$e^{2(m+n)}R^{-10} \sum_{\lambda_1,\dots,\lambda_{m+n}=1,2} \int dk_1 \cdots dk_{m+n}$$

 $\times I_{m+n,2,1^{m+n-1}} \left(e^2 \nu^2; |k_1|^2; \dots; |k_{m-1}|^2; |k_m|^2/R^2; |k_{m+1}|^2/R^2; |k_{m+2}|^2; \dots; |k_{m+n}|^2 \right)$
 $\times \left[\prod_{j \neq m, m+1} |k_j|^2 \hat{\varrho}^2(k_j) \right] \langle \varepsilon_1 | \varepsilon_2 \rangle \langle \varepsilon_2 | \varepsilon_3 \rangle \cdots \langle \varepsilon_{m+n} | \varepsilon_1 \rangle$
 $\times |k_m|^2 |k_{m+1}|^2 \hat{\varrho}^2(k_m/R) \hat{\varrho}^2(k_{m+1}/R) \cos(k_m \cdot \hat{n}) \cos(k_{m+1} \cdot \hat{n}).$ (B.10)

Switching to the polar coordinates as we did in the proof of Lemma 8.1, we see that

the r.h.s. of (B.10) =
$$e^{2(m+n)}R^{-9} \sum_{\lambda_1,\dots,\lambda_{m+n}=1,2} \left[\prod_{j=1}^{m+n} \int \mathrm{d}r_j \int \mathrm{d}X_j \int \mathrm{d}\varphi_j \right] \left[\frac{1}{\pi} \int_{-\infty}^{\infty} \mathrm{d}s \, s^2 \right]$$

 $\times \mathscr{F}_R(r_1,\dots,r_{m-1},r_{m+2},\dots,r_{m+n};s) \mathscr{G}_R(r_m,r_{m+1};s)$
 $\times \langle \varepsilon_1 | \varepsilon_2 \rangle \langle \varepsilon_2 | \varepsilon_3 \rangle \cdots \langle \varepsilon_{m+n} | \varepsilon_1 \rangle \cos(r_m X_m) \cos(r_{m+1} X_{m+1}),$ (B.11)

where

$$\mathscr{F}_{R} = \left(e^{2}\nu^{2} + R^{-2}s^{2}\right)^{-(m+n)} \left[\prod_{j \neq m, m+1} r_{j}^{4}\hat{\varrho}_{\mathrm{rad}}^{2}(r_{j})\left(r_{j}^{2} + R^{-2}s^{2}\right)^{-1}\right],$$
$$\mathscr{G}_{R} = r_{m}^{4}r_{m+1}^{4}\hat{\varrho}_{\mathrm{rad}}^{2}(r_{m}/R)\hat{\varrho}_{\mathrm{rad}}^{2}(r_{m+1}/R)\left(s^{2} + r_{m}^{2}\right)^{-1}\left(s^{2} + r_{m+1}^{2}\right)^{-1}.$$

Next, we will perform X_{j} - and φ_{j} -integrations for $j \neq m, m + 1$. For this purpose, we remark that

$$\sum_{\lambda=1,2} \int_{-1}^{1} \mathrm{d}X \int_{0}^{2\pi} \mathrm{d}\varphi \, |\varepsilon(k,\lambda)\rangle \langle \varepsilon(k,\lambda)| = \frac{4\pi}{3} \mathbb{1}_{3},$$

where $\mathbb{1}_3$ is the identity matrix acting in \mathbb{C}^3 . Using this and the fact that

$$\sum_{\lambda_m,\lambda_{m+1}=1,2} (\langle \varepsilon_m | \varepsilon_{m+1} \rangle_3)^2 = 1 + (\hat{k}_m \cdot \hat{k}_{m+1})^2,$$

we get

the r.h.s. of (B.11) =
$$e^{2(m+n)}R^{-9}\left(\frac{4\pi}{3}\right)^{m+n-2}\left[\frac{1}{\pi}\int_{-\infty}^{\infty}\mathrm{d}ss^2\right]\left[\prod_{j\neq m,m+1}\int\mathrm{d}r_j\right]\mathscr{F}_R$$

 $\times\left[\int\mathrm{d}r_m\mathrm{d}r_{m+1}\int\mathrm{d}X_m\mathrm{d}X_{m+1}\right]\mathfrak{S}(X_m,X_{m+1})\mathscr{G}_R\cos(r_mX_m)\cos(r_{m+1}X_{m+1}),$ (B.12)

where $\mathfrak{S}(X_m, X_{m+1})$ is defined by (B.5). Because

$$\mathscr{F}_R \le \left(e^2 \nu^2\right)^{-(m+n)} \left[\prod_{j \neq m, m+1} r_j^2 \hat{\varrho}_{\mathrm{rad}}^2(r_j)\right],$$

we obtain that

the r.h.s. of (B.12)
$$\leq R^{-9} \left(\frac{\|\hat{\varrho}\|_{L^2}^2}{3\nu^2} \right)^{m+n-2} \nu^{-4} \frac{1}{\pi} \int_{-\infty}^{\infty} \mathrm{d}s \, s^2 \int \mathrm{d}r_m \mathrm{d}r_{m+1} \\ \times \int \mathrm{d}X_m \mathrm{d}X_{m+1} \, \mathscr{G}_R \mathfrak{S}(X_m, X_{m+1}) \cos(r_m X_m) \cos(r_{m+1} X_{m+1}).$$
(B.13)

Here, we used the fact that the factor

$$\int \mathrm{d}r_m \int \mathrm{d}r_{m+1} \int \mathrm{d}X_m \mathrm{d}X_{m+1} [\cdots]$$

in the r.h.s. of (B.13) is positive for all $s \ge 0$. Using the elementary formula

$$\frac{1}{\pi} \int_{-\infty}^{\infty} \mathrm{d}s \, \frac{s^2}{\left(s^2 + r_m^2\right) \left(s^2 + r_{m+1}^2\right)} = \frac{1}{r_m + r_{m+1}} = \int_0^{\infty} \mathrm{d}t \, \mathrm{e}^{-t(r_m + r_{m+1})},$$

we have

the r.h.s. of (B.13) =
$$R^{-9} \left(\frac{\|\hat{\varrho}\|_{L^2}^2}{3\nu^2} \right)^{m+n-2}$$

 $\times \nu^{-4} \int_0^\infty dt \left\{ \frac{3}{2} \llbracket 1 \rrbracket \llbracket 1 \rrbracket - \llbracket 1 \rrbracket \llbracket [X^2] \rrbracket + \frac{3}{2} \llbracket X^2 \rrbracket \rrbracket \llbracket X^2 \rrbracket \right\},$ (B.14)

where

$$\llbracket A(X) \rrbracket(t) = (2\pi) \int_0^\infty \mathrm{d}r \, r^4 \mathrm{e}^{-tr} \hat{\varrho}_{\mathrm{rad}}^2(r/R) \int_{-1}^1 \mathrm{d}X \mathrm{e}^{\mathrm{i}rX} A(X).$$

We can compute $[\![1]\!]$ and $[\![X^2]\!]$ as

$$[[1]] = \int_{\mathbb{R}} \mathrm{d}x \, \frac{-24t^3 + 24t(1+x/R)^2}{\left\{t^2 + (1+x/R)^2\right\}^4} \varphi(x),$$
$$[[X^2]] = \int_{\mathbb{R}} \mathrm{d}x \, F_t(1+x/R)\varphi(x),$$

where

$$F_t(a) = 8 \frac{(1-a)t^5 - 2a(a^2 + a - 3)t^3 - a^3(a^2 + 3a + 6)t}{(t^2 + a^2)^4}.$$

Since $\varphi(x)$ decays rapidly, we readily see that the integral in (B.14) is uniformly bounded provided that R is sufficiently large.

C Basic properties of $I_{n_a,n_b,n_c}(a;b;c)$

Here, we will give a list of basic properties of $I_{n_a,n_b,n_c}(a;b;c)$ defined by (B.2).

The following result is easily checked:

$$I_{1,1,1}(a;b;c) = \frac{1}{ABC},$$
(C.1)

where $A = \sqrt{a} + \sqrt{b}$, $B = \sqrt{b} + \sqrt{c}$ and $C = \sqrt{c} + \sqrt{a}$. Using this, we have

$$I_{2,2,1}(a;b;c) = \frac{1}{4\sqrt{ab}}I_{1,1,1}(a;b;c)\left\{\frac{2}{A^2} + \frac{1}{AC} + \frac{1}{AB} + \frac{1}{BC}\right\},$$
(C.2)

$$I_{2,1,2}(a;b;c) = \frac{1}{4\sqrt{ac}} I_{1,1,1}(a;b;c) \left\{ \frac{2}{C^2} + \frac{1}{AC} + \frac{1}{BC} + \frac{1}{AB} \right\}$$
(C.3)

and

$$I_{3,1,1}(a;b;c) = \frac{1}{8a}I_{1,1,1}(a;b;c)\left\{\frac{2}{A^2} + \frac{2}{C^2} + \frac{2}{AC} + \frac{1}{\sqrt{aA}} + \frac{1}{\sqrt{aC}}\right\}.$$
(C.4)

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