

ON THE HYPERBOLIC DIRICHLET TO NEUMANN FUNCTIONAL

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Abstract: We prove the injectivity of the linearization of the hyperbolic Dirichlet to Neumann functional associated to metrics near the euclidean one in a “small” bounded domain of \mathbb{R}^3 , under some suitable transversality and geometric conditions.

1 – Introduction and statement of the results

Let \mathcal{M} denote the set of all riemannian metrics g on \mathbb{R}^n which coincide with the euclidean metric e , outside a bounded domain Ω with smooth boundary $\partial\Omega$. We consider the anisotropic wave equation

$$(1.1) \quad \begin{aligned} \square_g u &= \frac{\partial^2 u}{\partial t^2} - \Delta_g u = 0 \quad \text{in } \Omega \times (0, T) , \\ u &= f \quad \text{on } \Gamma = \partial\Omega \times (0, T), \quad f \in C_0^\infty(\Gamma) , \\ u &= \frac{\partial u}{\partial t} = 0 \quad \text{in } \Omega \times \{0\} . \end{aligned}$$

There is a unique solution to (1.1); hence we may define the hyperbolic Dirichlet to Neumann map as the linear operator

$$(1.2) \quad \Lambda_g: C_0^\infty(\Gamma) \rightarrow C^\infty(\Gamma) ,$$

$$(1.3) \quad \Lambda_g f = du \cdot \nu_g \Big|_\Gamma = \frac{\partial u}{\partial \nu_g} \Big|_\Gamma ,$$

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where u is the unique solution to (1.1) and ν_g is the g -outward unit normal to $\partial\Omega$. The hyperbolic Dirichlet to Neumann Functional:

$$(1.4) \quad \begin{aligned} \Lambda: \mathcal{M} &\rightarrow O_p(\Gamma) , \\ g &\mapsto \Lambda_g , \end{aligned}$$

where $O_p(\Gamma)$ denotes the space of all linear operators from $C_0^\infty(\Gamma)$ into $C^\infty(\Gamma)$, is known to be invariantly defined on the orbit obtained by the action over \mathcal{M} , of the group \mathcal{D} of all diffeomorphism ψ of $\bar{\Omega}$, each of which restricts to the identity on $\partial\Omega$. In fact, any such ψ can be used to construct a new metric, the pull-back metric, ψ^*g , such that $\Lambda_{\psi^*g} = \Lambda_g$. A natural conjecture is that this is the only obstruction to the uniqueness of Λ .

For fixed g , we consider the following map:

$$(1.5) \quad \psi \in \mathcal{D} \xrightarrow{A_g} \psi^*g \in \mathcal{M} .$$

It is easy to see that the tangent space $T_I\mathcal{D}$ of \mathcal{D} at the identity mapping I is the vector space $\Gamma_0(T\bar{\Omega})$ of all smooth vector fields on $\bar{\Omega}$ which vanish on $\partial\Omega$. On the other hand, the tangent space $T_g\mathcal{M}$ of \mathcal{M} at g is the vector space $\Gamma_0(S^2\bar{\Omega})$ of all smooth sections of symmetric 2-tensors on \mathbb{R}^n which are supported on $\bar{\Omega}$. We introduce respectively on $\Gamma_0(T\bar{\Omega})$ and on $\Gamma_0(S^2\bar{\Omega})$, the inner products

$$(1.6) \quad \langle X, Y \rangle = \int_{\bar{\Omega}} g(X, Y) v_g, \quad X, Y \in \Gamma_0(T\bar{\Omega}) ,$$

$$(1.7) \quad \langle\langle m, l \rangle\rangle = \frac{1}{n} \int_{\bar{\Omega}} \text{tr}(\hat{m} \circ \hat{l}) v_g, \quad m, l \in \Gamma_0(S^2\bar{\Omega}) ,$$

where v_g (resp. tr) denote the volume element (resp. the trace) associated to g and \hat{m} is the unique linear map (in fact a section of $\text{End}(T\bar{\Omega})$) defined by

$$(1.8) \quad g(\hat{m}u, v) = m(u, v), \quad \text{for all } u, v \in \Gamma(T\bar{\Omega}) .$$

Of course, \hat{g} is the identity on $\Gamma(T\bar{\Omega})$ and the factor $1/n$ in (1.7) is taken so as to have $\langle\langle g, g \rangle\rangle = \text{vol}_g(\bar{\Omega})$.

Consider as in [1], the formal linearizations of A_g at I and of Λ at g , respectively:

$$(1.9) \quad A'_g[I] := A'_g: \Gamma_0(T\bar{\Omega}) \rightarrow \Gamma_0(S^2\bar{\Omega})$$

and

$$(1.10) \quad \Lambda'_g: \Gamma_0(S^2\bar{\Omega}) \rightarrow O_p(\Gamma) .$$

Let $(A'_g)^*$ denote the formal adjoint of A'_g with respect to the inner product (1.6) and (1.7) and $\text{diam}_g(\Omega)$ the diameter of Ω in the metric g . In [1] the authors stated the following

Conjecture 1. *Let $m \in \Gamma_0(S^2\overline{\Omega})$ and assume that*

a) $\Lambda'_g(m) = 0,$

b) $(A'_g)^*(m) = 0$ and

c) $\text{diam}_g(\Omega) < T$ is sufficiently small that the exponential map for g is a global diffeomorphism in $\overline{\Omega}$.

Then m is identically zero.

Remark 1. The Condition b) in Conjecture 1.1 is obviously necessary. In fact, the range of A'_g is contained in the kernel of Λ'_g . Therefore, we should expect that Λ'_g be injective on a “transversal” subspace of the range of A'_g ; hence we shall refer to Condition b) as the Transversality Condition. The Condition c) is necessary to avoid the appearance of caustics.

Remark 2. Cardoso and Mendoza, [1], proved that Conjecture 1.1 holds if $n \geq 2$ and g is the euclidean metric e ; they also proved the conjecture when $n = 2$ and g is near the euclidean metric in the C^3 topology.

The main result of this paper is:

Theorem 1. *Conjecture 1.1 holds if $n = 3$, g is near the euclidean metric in the C^3 topology and in addition, one of the following two conditions is true:*

I – *The g Levi-Civita connection commute with rotation, i.e., $\nabla^g \circ J = J \circ \nabla^g$ (see Section 4 for the definition of J).*

II – *The generalized gradients of solutions of the eikonal equation are g -Killing fields (see [2] for the definition).*

The article is organized as follows: In Section 2 and 3 we develop the necessary preliminaries dealing with invariant formulas for A'_g , $(A'_g)^*$ and Λ'_g and the generalized X -ray and Radon transform. In Section 4 we present the proof of Theorem 1.1 with condition I and in Section 5 we prove Theorem 1.1 with condition II.

2 – Invariant formulas

Cardoso and Mendoza, [1], proved the following two propositions:

Proposition 1. *If $X \in \Gamma_0(T\bar{\Omega})$ and $m \in \Gamma_0(S^2\bar{\Omega})$, then it follows that*

$$(2.1) \quad A'_g(X)(\cdot, \cdot) = g(\nabla X, \cdot) + g(\cdot, \nabla X) ,$$

$$(2.2) \quad (A'_g)^*(m)(\cdot) = -\frac{2}{n} \sum_{i=1}^n \nabla_{e_i} m(\cdot, e_i) .$$

In (2.1) ∇ denotes the g Levi-Civita connection on $\Gamma_0(T\bar{\Omega})$ and in (2.2) ∇ is the g Levi-Civita connection on $\Gamma_0(S^2\bar{\Omega})$ and $(e_i)_{i=1,\dots,n} \in \Gamma(T\bar{\Omega})$, is a g orthonormal frame. We also observe that the right-hand side of (2.2) is independent of the chosen orthonormal frame.

We denote by $\tilde{m} \in \tilde{\Gamma}_0(S^2\bar{\Omega})$ the symmetric 2-tensor on $\Gamma(T^*\bar{\Omega})$ corresponding to m via g , i.e. $\tilde{m}(U^\#, V^\#) = m(U, V)$ for all $U, V \in \Gamma(T(\bar{\Omega}))$, where $U^\#(\cdot) = g(U, \cdot)$. We have the following:

Proposition 2. *The linearization of Λ at $g \in \mathcal{M}$, satisfies*

$$(2.3) \quad \langle \Lambda'_g(m)f, h \rangle_{L^2(\Gamma)} = \int_0^T \int_{\bar{\Omega}} \left\{ \tilde{m}(du, dv) + \frac{1}{2} \text{tr}(\hat{m}) [\tilde{g}(du, dv) - u_t v_t] \right\} v_g dt ,$$

for all $f, h \in C_0^\infty(\Gamma)$, where u is a solution of (1.1), v is a solution of

$$(2.4) \quad \square_g v = 0 \text{ in } \Omega \times (0, T), \quad v = v_t = 0 \text{ in } \Omega \times \{T\}, \quad v|_\Gamma = h ,$$

and $\langle \cdot, \cdot \rangle_{L^2(\Gamma)}$ is the L^2 -inner product in Γ with respect to the measure induced by the metric $g \otimes dt^2$.

3 – The geodesic X -ray transform and the Radon transform

Let $g \in \mathcal{M}$. We remind that g coincides with the euclidean metric, e , outside $\bar{\Omega}$. We shall deal with sections of the following vector bundles

$$\begin{array}{cc} \mathcal{P} & \mathcal{Q} \\ \downarrow & \downarrow \\ G & G \end{array}$$

where G denotes the manifold of geodesics (with respect to g), \mathcal{P} the bundle of parallel vector fields and Q the bundle of quadratic forms on \mathcal{P} . The generalized X -ray transform \mathcal{R}_g is the map

$$(3.1) \quad \mathcal{R}_g: \Gamma_0(S^2\overline{\Omega}) \rightarrow \Gamma(Q) ,$$

defined by

$$(3.2) \quad \mathcal{R}_g(m)_\gamma(P_1, P_2) = \int_\gamma m(\gamma(t)) \left(P_1(\gamma(t)), P_2(\gamma(t)) \right) ,$$

where $\gamma \in G$ and $P_1, P_2 \in \Gamma(\mathcal{P})$. There is a global non-vanishing section $T: G \rightarrow \mathcal{P}$, given by

$$(3.3) \quad T_\gamma(\gamma(t)) = \dot{\gamma}(t) ,$$

since as it is well known if γ is a geodesic, then $\dot{\gamma}(t) = P_{\gamma,0,t}(\dot{\gamma}(0))$, where $P_{\gamma,0,t}$ is the parallel transport along γ , from 0 to t . In [1] it was proved the following:

Proposition 3. *Let $m \in \Gamma_0(S^2\overline{\Omega})$ satisfy $\Lambda'_g(m) = 0$ and $\text{diam}_g(\Omega) < T$ be so small that the exponential map for g is a global diffeomorphism in $\overline{\Omega}$. Then*

$$(3.4) \quad \mathcal{R}_g(m)_\gamma(T_\gamma, T_\gamma) = 0 ,$$

for all g -geodesic γ .

Let \mathcal{G}' denote the space of generalized hiperplanes $\Sigma = \Sigma_\phi^s = \Sigma(\phi, \eta, s)$, where $s \in \mathbb{R}$, $\eta \in S^{n-1}$ is a normal vector to Σ and $\phi(\cdot, \eta)$ is a solution of the eikonal equation

$$(3.5) \quad \begin{cases} g(\nabla^g \phi(\cdot, \eta), \nabla^g \phi(\cdot, \eta)) = 1 , \\ \phi(\cdot, \eta)|_\Sigma = s , \\ \nabla^g \phi(\cdot, \eta)|_\Sigma = \eta . \end{cases}$$

We assume that

- (i) the metric g satisfies Condition II of Theorem 1.1.
- (ii) $\phi(x, tw) = t\phi(x, w)$, for all $(w, t) \in S^{n-1} \times \mathbb{R}$.

Remark 3. The generalized hiperplanes are closed submanifolds of dimension $n-1$. On the other hand taking into account (i) it is easy to see that they are totally geodesic submanifolds.

Let Q' denote the quadratic bundle over \mathcal{G}' . The generalized Radon transform R_g is the map

$$(3.6) \quad R_g: \Gamma_0(S^2\overline{\Omega}) \rightarrow \Gamma(Q') ,$$

defined by

$$(3.7) \quad R_g(m)_\Sigma(X, Y) = \int_\Sigma m(X, Y) \mu_\Sigma ,$$

where μ_Σ denotes the volume element induced on Σ by the metric g .

Corollary 1. *There is an orthonormal frame T_1, \dots, T_{n-1}, N of $T\mathbb{R}^n$ such that*

$$(3.8) \quad R_g(m)_\Sigma(T_i, T_j) = 0 ,$$

for all $i, j \in \{1, \dots, n-1\}$.

Proof: Let $\Sigma = \Sigma_\phi^\lambda$ in \mathcal{G}' and $\phi_i \in C^\infty(\mathbb{R}^n)$, $i = 1, \dots, n-1$, such that

$$(3.9) \quad g(\nabla^g \phi, \nabla^g \phi_i) = 0 ,$$

$$(3.10) \quad g(\nabla^g \phi_i, \nabla^g \phi_j) = \delta_{ij} .$$

Denoting $T_i := \nabla^g \phi_i$ and $N := \nabla^g \phi$, $i = 1, \dots, n-1$, it follows from (3.9) and (3.10) that $T_i|_\Sigma \in \Gamma(T\Sigma)$ and T_1, \dots, T_{n-1}, N is an orthonormal frame of $T\mathbb{R}^n$.

We can assume that $\Sigma \cap \Sigma_{\phi_i}^0 \cap \Sigma_{\phi_j}^0$ is not-empty for all $i, j \in \{1, \dots, n-1\}$ and denote

$$\mathcal{N}_i := \Sigma \cap \Sigma_{\phi_i}^0 ,$$

$$\mathcal{N}_{ij} := \Sigma \cap \Sigma_{\phi_i}^0 \cap \Sigma_{\phi_j}^0 .$$

Let Φ_i^σ be the geodesic flow associated to the field $\nabla^g \phi_i$, then $\sigma \mapsto \Phi_i^\sigma(\cdot)$ are the geodesics which start at \mathcal{N}_i and, using (3.4), we obtain

$$\begin{aligned} R_g(m)_\Sigma(T_i, T_j) &= \int_{-\infty}^\infty \int_{\mathcal{N}_i} m(\Phi_i^\sigma(y)) \left(\dot{\Phi}_i^\sigma(y), \dot{\Phi}_i^\sigma(y) \right) dS_y d\sigma \\ &= \int_{\mathcal{N}_i} \mathcal{R}_g(m)_{\Phi_i^\bullet(y)} \left(\dot{\Phi}_i^\bullet(y), \dot{\Phi}_i^\bullet(y) \right) dS_y \\ &= 0 . \end{aligned}$$

A similar calculation holds for Φ_{ij}^σ , the geodesic flow associated to the field $\nabla^g \xi_{ij}$, where $\xi_{ij} = (\phi_i + \phi_j)/\sqrt{2}$, taking into account that m is symmetric and (3.4). We obtain

$$R_g(m)_\Sigma(T_i, T_j) = \frac{1}{2} R_g(m)_\Sigma(T_i + T_j, T_i + T_j) = 0 ,$$

which concludes the proof. ■

4 – Proof of Theorem 1.1 with condition I

In this section $\bar{\Omega}$ will be a smooth bounded domain in \mathbb{R}^3 . It is convenient that $\bar{\Omega}$ be placed in the open first octant in \mathbb{R}^3 . Consider the vector bundles \mathcal{P} and Q over G as in the beginning of Section 3. A section $m = (m_{ij})_{i,j=1,2,3} \in \Gamma_0(S^2\bar{\Omega})$ belongs to $L^2(S^2\bar{\Omega})$ if

$$\|m\|_0^2 = \int_{\bar{\Omega}} \left(m_{11}^2 + m_{22}^2 + m_{33}^2 + 2(m_{13}^2 + m_{12}^2 + m_{23}^2) \right) dx_1 dx_2 dx_3 < \infty ,$$

where

$$m_{ij} = m \left(\frac{\partial}{\partial x_i}, \frac{\partial}{\partial x_j} \right) ,$$

and x_1, x_2 and x_3 are the standard euclidean coordinates in \mathbb{R}^3 . The corresponding Sobolev space based on $L^2(S^2\bar{\Omega})$ will be denoted by $H^s(S^2\bar{\Omega})$.

There is a natural frame for \mathcal{P} , manely

$$(4.1) \quad T = \dot{\gamma}, \quad N = J\dot{\gamma}, \quad M ,$$

where $\dot{\gamma}$ is the unit tangent vector to the g -geodesic γ , J denotes the $\frac{\pi}{2}$ clockwise rotation (with respect to g) in the plane generated by T and the axis oz and M is the parallel transport along γ of the vector product of the euclidean counterparts of T and N . We say that a section H of Q is in $L^2(Q)$ if

$$\|H\|_0^2 = \int_G \left(H_{TT}^2 + H_{NN}^2 + H_{MM}^2 + 2(H_{TN}^2 + H_{TM}^2 + H_{NM}^2) \right) d\mu < \infty ,$$

where $H_{AB} := H(A, B)$, and $d\mu$ represents the naturally defined Liouville measure on G . The corresponding Sobolev space will be denoted by $H^s(Q)$. We shall need the following (see [5]):

Lemma 1. *If g is near the euclidean metric in the C^3 topology, then*

$$\mathcal{R}_g: H_{\text{comp}}^s(S^2\bar{\Omega}) \rightarrow H_{\text{loc}}^{s+1/2}(Q)$$

is a bounded linear operator with a bounded inverse.

We shall also introduce local coordinates in G , parametrizing a geodesic by $(x_1, x_2, \theta, \varphi)$ or $(x_2, x_3, \theta, \varphi)$ or $(x_1, x_3, \theta, \varphi)$, where (x_1, x_2) (resp. (x_2, x_3) , resp. (x_1, x_3)) is the point of intersection of the geodesic with the x_1x_2 (resp. x_2x_3 , resp. x_1x_3) plane and (θ, φ) is the spacial position of the speed vector. Let

$$(4.2) \quad T = T(x_1, x_2, \theta, \varphi, t), \quad N = N(x_1, x_2, \theta, \varphi, t), \quad M = M(x_1, x_2, \theta, \varphi, t) ,$$

be the orthonormal frame with respect to g , defined by (4.1). We assume that

$$(4.3) \quad t \rightarrow M(x_1, x_2, \theta, \varphi, t) \quad \text{and} \quad t \rightarrow N(x_1, x_2, \theta, \varphi, t)$$

are extended as odd functions for $t \leq 0$. We also denote

$$(4.4) \quad \begin{aligned} \Theta_1 &:= \gamma_* \left(\frac{\partial}{\partial \theta} \right) = \beta_0 N, \\ \Theta_2 &:= \gamma_* \left(\frac{\partial}{\partial \varphi} \right) = \delta_0 M, \\ X_i &:= \gamma_* \left(\frac{\partial}{\partial x} \right) = \alpha_i T + \beta_i N + \delta_i M, \quad i = 1, 2, \end{aligned}$$

where β_j and δ_j , $j = 0, 1, 2$ are functions that depend of the variables $x_1, x_2, \theta, \varphi, t$ and α_1, α_2 only depend on the variables x_1, x_2, θ and φ . To see this, we note that

$$Tg(\Theta_i, T) = g(\nabla_T \Theta_i, T) = g(\nabla_{\Theta_i} T, T) = \frac{1}{2} \Theta_i g(T, T) = 0,$$

and

$$Tg(X_i, T) = g(\nabla_T X_i, T) = g(\nabla_{X_i} T, T) = \frac{1}{2} X_i g(T, T) = 0.$$

Now, at $t=0$,

$$\Theta_1 \Big|_{t=0} = \frac{\partial \dot{\gamma}}{\partial \theta} = -N,$$

and

$$\Theta_2 \Big|_{t=0} = \frac{\partial \dot{\gamma}}{\partial \varphi} = \cos \theta M.$$

If g is the euclidean metric, geodesics are straight lines and in this case $T = \cos \theta e^{i\varphi} + \sin \theta e_3$, $N = \sin e^{i\varphi} - \cos \theta e_3$ and $M = i e^{i\varphi}$, where $(e_i)_{i=1,2,3}$ is the canonical basis in \mathbb{R}^3 .

It is easy to see that when g is nearly euclidean i.e. $\|g - e\|_{C^k(\bar{\Omega})} \leq \delta$, then T, N, M, Θ_i and X_i are close to their euclidean counterparts, so that we may assume that

$$(4.5) \quad \sup_{\substack{0 \leq x, y, t \leq L \\ 0 \leq \theta, \varphi \leq \pi/4}} \left\{ \sum_{i=0}^2 \left(\left| \frac{\partial \beta_i}{\partial t} \right| + \left| \frac{\partial \delta_i}{\partial t} \right| \right) + |N(\beta_0)| + |M(\delta_0)| \right\} \leq \varepsilon,$$

where L is the length of the sides of the isosceles triangles with sides on the coordinates axis whose faces generate a prisme which completely encloses $\bar{\Omega}$. We shall need the

Lemma 2. *The following identities hold:*

$$(4.6) \quad \nabla_{\Theta_1} M = \nabla_N M = 0 ,$$

$$(4.7) \quad \nabla_{\Theta_1} N = -T(\beta_0) T; \quad \nabla_N N = -\frac{T(\beta_0)}{\beta_0} T ,$$

$$(4.8) \quad \nabla_{\Theta_2} N = T(\delta_0) M; \quad \nabla_M N = \frac{T(\delta_0)}{\delta_0} M ,$$

$$(4.9) \quad \nabla_{\Theta_2} M = -T(\delta_0) (T + N); \quad \nabla_M M = -\frac{T(\delta_0)}{\delta_0} (T + N) ,$$

$$(4.10) \quad \nabla_M T = \frac{T(\delta_0)}{\delta_0} M; \quad \nabla_N T = \frac{T(\beta_0)}{\beta_0} N ,$$

$$(4.11) \quad \nabla_{X_i} N = -T(\beta_i) T + T(\delta_i) M; \quad i = 1, 2 ,$$

$$(4.12) \quad \nabla_{X_i} M = -T(\delta_i) (T + N); \quad i = 1, 2 ,$$

$$(4.13) \quad \frac{T(\delta_i)}{\delta_i} = \frac{T(\delta_j)}{\delta_j}; \quad \frac{T(\beta_i)}{\beta_i} = \frac{T(\beta_j)}{\beta_j}; \quad i, j = 0, 1, 2 ,$$

$$(4.14) \quad M(\beta_i) = 0; \quad T(\delta_i) = N(\delta_i); \quad i = 0, 1, 2 ,$$

$$(4.15) \quad N(\alpha_i) = M(\alpha_i) = 0; \quad i = 1, 2 ,$$

$$(4.16) \quad \frac{N(\beta_i)}{\beta_i} = \frac{N(\beta_j)}{\beta_j}; \quad \frac{M(\delta_i)}{\delta_i} = \frac{M(\delta_j)}{\delta_j}; \quad i, j = 0, 1, 2 .$$

Proof: Because (taking into account Condition I)

$$\nabla_{\Theta_1} N = \nabla_{\Theta_1} J T = J \nabla_{\Theta_1} T = J \nabla_T \Theta_1 = J T(\beta_0) N = -T(\beta_0) T ,$$

we obtain (4.7).

To establish (4.12), write

$$\nabla_{X_i} M = a T + b N ,$$

where we have

$$\begin{aligned} a &= -g(M, \nabla_T X_i) \\ &= -T(\beta_i) g(M, N) - T(\delta_i) g(M, M) \\ &= -T(\delta_i) , \end{aligned}$$

and

$$\begin{aligned}
b &= -g(M, \nabla_{X_i} N) \\
&= T(\beta_i) g(M, T) - T(\delta_i) g(M, M) \\
&= -T(\delta_i) .
\end{aligned}$$

To establish (4.14), (4.15) and (4.26), we note that $[\Theta_1, \Theta_2] = 0$; in this way,

$$\begin{aligned}
\nabla_N M - \nabla_M N &= [N, M] \\
&= \frac{M(\beta_0)}{\beta_0} N - \frac{N(\delta_0)}{\delta_0} M .
\end{aligned}$$

Using (4.6) we obtain

$$\nabla_M N = \frac{N(\delta_0)}{\delta_0} M - \frac{M(\beta_0)}{\beta_0} N ,$$

and if we compare with (4.8), we conclude that $M(\beta_0) = 0$ and $T(\delta_0) = N(\delta_0)$.

On the other hand, using the fact that $[X_i, \Theta_1] = 0$, we obtain

$$\nabla_{X_i} N - \nabla_N X_i = [X_i, N] = -\frac{X_i(\beta_0)}{\beta_0} N ;$$

thus

$$\begin{aligned}
\nabla_{X_i} N &= -\frac{X_i(\beta_0)}{\beta_0} N + \nabla_N X_i \\
&= \left(-\frac{X_i(\beta_0)}{\beta_0} + \alpha_i \frac{T(\beta_0)}{\beta_0} + N(\beta_i) \right) N + \left(N(\alpha_i) - \beta_i \frac{T(\beta_0)}{\beta_0} \right) T + N(\delta_i) M \\
&= \left(N(\beta_i) - \beta_i \frac{N(\beta_0)}{\beta_0} \right) N + \left(N(\alpha_i) - T(\beta_i) \right) T + N(\delta_i) M .
\end{aligned}$$

Now comparing with (4.11), we obtain

$$N(\alpha_i) = 0, \quad T(\delta_i) = N(\delta_i), \quad \frac{N(\beta_i)}{\beta_i} = \frac{N(\beta_0)}{\beta_0}, \quad i = 1, 2 .$$

It follows from similar computations (taking into account that $[X_i, \Theta_2] = 0$) that

$$M(\beta_i) = M(\alpha_i) = 0, \quad \frac{M(\delta_i)}{\delta_i} = \frac{M(\delta_0)}{\delta_0}, \quad i = 1, 2 ;$$

this concludes the proof. ■

We introduce the following notation:

$$A_i := \frac{\beta_i}{\delta_0} T(\delta_0), \quad B_i := \frac{\delta_i}{\delta_0} M(\delta_0), \quad C_i := \frac{\delta_i}{\beta_0} T(\delta_0), \quad D_i := \frac{\beta_i}{\beta_0} N(\beta_0) .$$

We shall need the

Proposition 4. *If m satisfies the hypotheses of Theorem 1.1 with Condition I and $H = \mathcal{R}_g(m)$, then the following system of equations holds:*

$$(4.17) \quad \Theta_i(H_{TT}) = X_i(H_{TT}) = 0, \quad i = 1, 2,$$

$$(4.18) \quad \Theta_1(H_{NN}) = - \int_{\gamma} T(\beta_0) m(N, T) + A_0(m(N, T+N) - m(M, M)),$$

$$(4.19) \quad \Theta_2(H_{NN}) = - \int_{\gamma} M(\delta_0) m(N, N),$$

$$(4.20) \quad X_i(H_{NN}) = - \int_{\gamma} T(\beta_i) m(N, T) + A_i(m(N, T+N) - m(M, M)) \\ + B_i m(N, N),$$

$$(4.21) \quad \Theta_1(H_{MM}) = - \int_{\gamma} N(\beta_0) m(M, M),$$

$$(4.22) \quad \Theta_2(H_{MM}) = - \int_{\gamma} T(\delta_0) m(M, T+N) + C_0 m(N, T),$$

$$(4.23) \quad X_i(H_{MM}) = - \int_{\gamma} T(\delta_i) m(M, T+N) + C_i m(M, T) + D_i m(M, M),$$

$$(4.24) \quad \Theta_1(H_{TN}) = \int_{\gamma} T(\beta_0) m(N, N) + A_0(m(T, T+N) - m(M, M)),$$

$$(4.25) \quad \Theta_2(H_{TN}) = - \int_{\gamma} M(\delta_0) m(T, N),$$

$$(4.26) \quad X_i(H_{TN}) = \int_{\gamma} T(\beta_i) m(N, N) - A_i(m(T, T+N) - m(M, M)) \\ - B_i m(T, N),$$

$$(4.27) \quad \Theta_1(H_{NM}) = 2 \int_{\gamma} A_0 m(T+N, M),$$

$$(4.28) \quad \Theta_2(H_{NM}) = \int_{\gamma} T(\delta_0) m(M, M) - 2 C_0 m(T, N),$$

$$(4.29) \quad X_i(H_{NM}) = \int_{\gamma} T(\delta_i) m(M, M) - 2 C_i m(T, N) - D_i m(N, M),$$

$$(4.30) \quad \Theta_1(H_{TM}) = - \int_{\gamma} N(\beta_0) m(T, M) ,$$

$$(4.31) \quad \Theta_2(H_{TM}) = \int_{\gamma} T(\delta_0) m(M, M) - C_0(m(T, T) - m(N, N)) ,$$

$$(4.32) \quad X_i(H_{TM}) = \int_{\gamma} T(\delta_i) m(M, M) - C_i(m(T, T) - m(N, N)) \\ - D_i m(T, N) .$$

Proof: The meaning of (3.4) in Proposition 3.1 is that $H_{TT} = 0$, and hence (4.17) holds. The Transversality Condition satisfied by m means that

$$(4.33) \quad \nabla_T m(T, T) + \nabla_N m(T, N) + \nabla_M m(T, M) = 0 ,$$

$$(4.34) \quad \nabla_T m(N, T) + \nabla_N m(N, N) + \nabla_M m(N, M) = 0 ,$$

$$(4.35) \quad \nabla_T m(M, T) + \nabla_N m(M, N) + \nabla_M m(M, M) = 0 .$$

We begin by computing (4.18). Since $\nabla_{\Theta_1} N = -T(\beta_0) T$, using (4.34), after integrating by parts (taking into account (4.3), (4.8), (4.9), (4.15) and that m is compactly supported in $\bar{\Omega}$), we obtain

$$\begin{aligned} \Theta_1(H_{NN}) &= \int_{-\infty}^{\infty} \Theta_1 m(N, N) \\ &= \int_{-\infty}^{\infty} \nabla_{\Theta_1} m(N, N) + 2m(\nabla_{\Theta_1} N, N) \\ &= \int_{-\infty}^{\infty} -\beta_0 T m(T, N) - 2T(\beta_0) m(T, N) - \beta_0 M m(N, N) \\ &\quad + \int_{-\infty}^{\infty} \beta_0 m(\nabla_M M, N) + \beta_0 m(M, \nabla_M N) \\ &= \int_{-\infty}^{\infty} -T(\beta_0) m(T, N) + M(\beta_0) m(N, N) + A_0(m(M, M) - m(T+N, N)) \\ &= - \int_{-\infty}^{\infty} T(\beta_0) m(T, N) + A_0(m(T+N, N) - m(M, M)) . \end{aligned}$$

To establish (4.20) we proceed in a similar way. In fact,

$$X_i(H_{NN}) = \int_{-\infty}^{\infty} \beta_i \nabla_N m(N, N) + \delta_i \nabla_M m(N, N) + 2m(\nabla_{X_i} N, N) =$$

$$\begin{aligned}
 &= \int_{-\infty}^{\infty} -\beta_i Tm(N, T) - \beta_i Mm(N, M) + \delta_i Mm(N, N) - 2T(\beta_i) m(T, N) \\
 &\quad + \int_{-\infty}^{\infty} -2\delta_i m(\nabla_M N, N) + 2T(\delta_i) m(M, N) \\
 &\quad + \int_{-\infty}^{\infty} \beta_i m(\nabla_M N, M) + \beta_i m(N, \nabla_M M) \\
 &= \int_{-\infty}^{\infty} -T(\beta_i) m(T, N) + M(\beta_i) m(N, M) - M(\delta_i) m(N, N) \\
 &\quad + \int_{-\infty}^{\infty} A_i(m(M, M) - m(N, T+N)) \\
 &= - \int_{-\infty}^{\infty} T(\beta_i) m(T, N) + B_i m(N, N) + A_i(m(N, T+N) - m(M, M)) .
 \end{aligned}$$

The remaining equations follow from analogous computations. ■

Proof of Theorem 1.1 with Condition I: It follows from Proposition 4.1 and (4.5) that there is a constant $C_1 > 0$ such that

$$(4.36) \quad \|dH\|_{L^2(Q)} \leq C_1 \varepsilon \|m\|_{L^2(S^2\bar{\Omega})} ,$$

where ε can be made arbitrarily small by requiring that g be sufficiently close to the euclidean metric. On the other hand, by the Poincaré inequality, we obtain that there is a constant $C_2 > 0$ such that

$$(4.37) \quad \|H\|_{H^1(Q)} \leq C_2 \|dH\|_{L^2(Q)} ,$$

and using Lemma 4.1, it follows that there is a constant $C_3 > 0$ such that

$$(4.38) \quad \|m\|_{L^2(S^2\bar{\Omega})} \leq C_3 \|H\|_{H^{1/2}(Q)} \leq C_3 \|H\|_{H^1(Q)} .$$

Using (4.36)–(4.38) we get $H = 0$ if ε is chosen small enough and, consequently, $m = 0$. ■

5 – Proof of Theorem 1.1 with condition II

In this section $\bar{\Omega}$ will be a smooth domain in \mathbb{R}^3 . It is convenient that $\bar{\Omega}$ be placed in the open first octant in \mathbb{R}^3 . Consider the quadratic bundle Q' over \mathcal{G}' as in the beginning of Section 3. A section M of Q' belongs to $L^2(Q')$ if

$$(5.1) \quad \|M\|_0^2 = \int_{\mathcal{G}'} (M_{T_1 T_1}^2 + M_{T_2 T_2}^2 + M_{NN}^2 + 2(M_{T_1 N}^2 + M_{T_2 N}^2 + M_{T_1 T_2}^2)) d\mu < \infty ,$$

where $M_{AB} := M(A, B)$, $d\mu$ represents the naturally defined Liouville measure on \mathcal{G}' and T_1, T_2, N are the vector fields given by Corollary 3.1. The corresponding Sobolev space based on $L^2(Q')$ will be denoted by $H^s(Q')$. We shall need the

Lemma 3. *If g is near the euclidean metric in the C^3 topology, then*

$$R_g : H_{\text{comp}}^s(S^2\overline{\Omega}) \rightarrow H_{\text{loc}}^{s+\frac{n-1}{2}}(Q')$$

is a bounded linear operator with a bounded inverse.

Proof: The adjoint, R_g^* , of R_g is given by

$$R_g^*h(x) = \int_{S^{n-1}} h(\omega, \phi(x, \omega)) \, d\omega .$$

Let $P = (2\pi)^{1-n}R_g^* \partial_s^{n-1}R_g$. Using Fourier inversion formula, making $t\omega = \xi$ and observing that $dt \, d\omega = |\xi|^{1-n}d\xi$, we obtain

$$(5.2) \quad Pf(x) = (2\pi)^{-n} \int_{\mathbb{R}^n} \int_{\mathbb{R}^n} e^{i(\phi(x,\xi)-\phi(y,\xi))} f(y) \, d\xi \, dy .$$

By Taylor's formula with integral remainder, we get

$$(5.3) \quad \phi(x, \xi) - \phi(y, \xi) = \langle x - y, k_{x,y}(\xi) \rangle ,$$

where

$$(5.4) \quad k_{x,y}(\xi) = \int_0^1 D_x \phi(y + s(x-y), \xi) \, ds .$$

We observe that since g is near the euclidean metric, the function $k_{x,y}$ is a global diffeomorphism. If we substitute (5.3) in (5.2) we obtain that P is a pseudodifferential operator of order zero, with amplitude function given by

$$(5.5) \quad a(x, y, \xi) = \frac{1}{\det [D_\xi k_{x,y}(k_{x,y}^{-1}(\xi))]} .$$

Hence, we obtain from the standard estimates for pseudodifferential operators (see [4], Proposition 9.2) and the fact that $D_\xi k_{x,y}(\xi)$ is near the identity, that $\|P - I\|_{\mathcal{L}(L^2)} < 1$ and, consequently, R_g is invertible. ■

Let us consider the following open set

$$\mathcal{U} = \left\{ \Sigma \in \mathcal{G}' : \Sigma \text{ is transversal to the } x\text{-axis} \right\} .$$

We may parametrize a generalized hiperplane $\Sigma \in \mathcal{U}$ by $\Sigma = \Sigma(x, \theta, \varphi)$, where $x\mathbf{e}_1$ is the point of intersection of Σ with the x -axis and $N^e = N^e(\theta, \varphi)$ is the representation of the normal vector of Σ in spherical coordinates. We remind that the generalized hiperplanes are totally geodesic by Condition II. We consider the following map

$$\text{Exp}: \mathbb{R}^2 \rightarrow \Sigma ,$$

given by

$$\text{Exp}(x_1, x_2) = \text{Exp}_{x\mathbf{e}_1}(x_1T_1^e + x_2T_2^e) ,$$

where $T_1^e = \frac{\partial N^e}{\partial \theta}$, $T_2^e = \frac{\partial N^e}{\partial \varphi}$ and $\text{Exp}_{x\mathbf{e}_1}$ denote the exponential map at $x\mathbf{e}_1$. We write

$$M = R_g(m), \quad (\text{Exp})^*(\mu_\Sigma) = \delta dx_1 dx_2 .$$

Here μ_Σ is the volume element on Σ induced by the metric g . We may finally write:

$$M(X, Y) = \int_{\mathbb{R}^2} m(\text{Exp}(x)) (X \circ \text{Exp}(x), Y \circ \text{Exp}(x)) \delta(x) dx_1 dx_2 .$$

Let $\gamma(t, x, \theta, \varphi) = \text{Exp}(t(x_1, x_2))$ be the g -geodesic through x with tangent vector $x_1T_1^e + x_2T_2^e$, and consider the g -orthonormal fields

$$N = \nabla^g \phi, \quad T_1 = \nabla^g \phi_1, \quad T_2 = \nabla^g \phi_2 ,$$

given by Corollary 3.1, along the g -geodesic γ . Let

$$(5.6) \quad \begin{aligned} N &:= N(t, x, \theta, \varphi) = N(\gamma(t)), & T_1 &:= T_1(t, x, \theta, \varphi) = T_1(\gamma(t)) , \\ T_2 &:= T_2(t, x, \theta, \varphi) = T_2(\gamma(t)), & T &:= T(t, x, \theta, \varphi) = \dot{\gamma}(t) . \end{aligned}$$

We note that

$$(5.7) \quad \begin{aligned} \Theta_1 &:= \gamma_* \left(\frac{\partial}{\partial \theta} \right) = \beta_0 N , \\ \Theta_2 &:= \gamma_* \left(\frac{\partial}{\partial \varphi} \right) = \beta_1 T_1 + \beta_2 T_2 + \beta_3 N , \\ X &:= \gamma_* \left(\frac{\partial}{\partial x} \right) = \alpha_1 T_1 + \alpha_2 T_2 + \alpha_3 N , \end{aligned}$$

where the functions (see Lemma 5.2) $\alpha_1, \alpha_2, \beta_0, \beta_1$ and β_2 depend on the variables x, θ, φ and α_3, β_3 depend on the variables t, x, θ and φ .

Remark 4. If g is near the euclidean metric then T_1, T_2, N, X, δ and Θ_i are close to their euclidean counterparts.

We note that $T = x_1 T_1 + x_2 T_2$. In fact, since Σ is totally geodesic, it follows that N and T are g -orthonormal. Therefore $T = a_1 T_1 + a_2 T_2$; now using the fact that T_i is a g -Killing field, we obtain that a_i is constant.

We introduce the following notation

$$A := \frac{1}{2(\beta_1 x_2 - \beta_2 x_1)} .$$

Lemma 4. *The following statements hold:*

$$(5.8) \quad \nabla_N N = \nabla_{T_i} T_i = \nabla_{\Theta_i} N = 0, \quad i = 1, 2 ,$$

$$(5.9) \quad [N, T_1] = [N, T_2] = 0 .$$

$$(5.10) \quad \text{The coefficients } \alpha_1, \alpha_2, \beta_0, \beta_1, \beta_2 \text{ in (5.7)} \\ \text{are independent of the variable } t .$$

$$(5.11) \quad \nabla_{T_1} T_2 = AT(\beta_3) N, \quad \nabla_{T_2} T_1 = -AT(\beta_3) N ,$$

$$(5.12) \quad \nabla_{T_1} N = \nabla_N T_1 = -AT(\beta_3) T_2, \quad \nabla_{T_2} N = \nabla_N T_2 = AT(\beta_3) T_1 ,$$

$$(5.13) \quad \nabla_{\Theta_2} T_1 = -\beta_2 AT(\beta_3) N - \beta_3 AT(\beta_3) T_2 ,$$

$$(5.14) \quad \nabla_{\Theta_2} T_2 = \beta_1 AT(\beta_3) N + \beta_3 AT(\beta_3) T_1 ,$$

$$(5.15) \quad \nabla_{\Theta_2} N = -\beta_1 AT(\beta_3) T_2 + \beta_2 AT(\beta_3) T_1 ,$$

$$(5.16) \quad \nabla_{\Theta_1} T_1 = -\beta_0 AT(\beta_3) T_2, \quad \nabla_{\Theta_1} T_2 = \beta_0 AT(\beta_3) T_1 ,$$

$$(5.17) \quad N(\beta_1) = N(\beta_2) = N(\alpha_1) = N(\alpha_2) = 0 .$$

Proof: To establish (5.8), we write

$$\nabla_N N = a_1 T_1 + a_2 T_2 ;$$

since T_i is a g -Killing field, we have

$$a_i = g(\nabla_N N, T_i) = -g(N, \nabla_N T_i) = 0 .$$

The statement (5.9) is a consequence of $[\Theta_1, T] = 0$. To get (5.10), we observe that

$$(5.18) \quad \nabla_{T_1} T_2 = g(\nabla_{T_1} T_2, N) N, \quad \nabla_{T_2} T_1 = g(\nabla_{T_2} T_1, N) N .$$

Using (5.18) and the fact that $\nabla_{\Theta_2} T = \nabla_T \Theta_2$, we obtain that β_1 and β_2 are independent of the variable t . Futhermore,

$$(5.19) \quad 2 AT(\beta_3) = g(\nabla_{T_1} T_2, N) - g(\nabla_{T_2} T_1, N) .$$

Since $[X, T] = 0$, it follows that α_1 and α_2 are also independent of the variable t . We observe that

$$(5.20) \quad g(\nabla_{T_1} T_2, N) = -g(T_2, \nabla_{T_1} N) = -g(T_2, \nabla_N T_1) = -g(N, \nabla_{T_2} T_1) .$$

Using (5.18)–(5.20) we obtain (5.11). To establish (5.17), we use the facts that $[\Theta_1, \Theta_2] = [X, \Theta_1] = 0$. The remaining formulas are immediate. ■

We introduce the following notation:

$$m_{ij} := m(T_i, T_j), \quad m_{i0} := m(N, T_i), \quad m_{00} := m(N, N), \quad i, j = 1, 2 .$$

We shall need the

Proposition 5. *If m satisfies the hypotheses of Theorem 1.1 with Condition II and $M = R_g(m)$, then the following system of equations holds:*

$$(5.21) \quad X(M_{T_i T_j}) = \Theta_k(M_{T_i T_j}) = 0, \quad i, j, k = 1, 2 ,$$

$$(5.22) \quad \begin{aligned} X(M_{T_j N}) &= \sum_{i=1}^2 \int \left(T_i(\delta\alpha_3) m_{ij} - T_i(\delta\alpha_i) m_{jo} + (-1)^j \delta\alpha_3 AT(\beta_3) m_{io} \right) \\ &\quad + \int \left((-1)^j \delta\alpha_3 AT(\beta_3) m_{k(j)o} + X(\delta) m_{jo} \right) , \end{aligned}$$

$$(5.23) \quad \begin{aligned} \Theta_1(M_{T_j N}) &= \sum_{i=1}^2 \int \left(T_i(\delta\beta_0) m_{ij} + (-1)^j \delta\beta_0 AT(\beta_3) m_{io} \right) \\ &\quad + \int \left((-1)^j \delta\beta_0 AT(\beta_3) m_{k(j)o} + \Theta_1(\delta) m_{jo} \right) , \end{aligned}$$

$$(5.24) \quad \begin{aligned} \Theta_2(M_{T_j N}) &= \sum_{i=1}^2 \int \left(T_i(\delta\beta_3) m_{ij} - T_i(\delta\beta_i) m_{jo} + (-1)^j \delta\beta_3 AT(\beta_3) m_{io} \right) \\ &\quad + \int \left((-1)^j \delta\beta_3 AT(\beta_3) m_{k(j)o} + \Theta_2(\delta) m_{jo} \right) , \end{aligned}$$

$$(5.25) \quad X(M_{NN}) = \sum_{i=1}^2 \int \left(T_i(\delta\alpha_3) m_{io} - T_i(\delta\alpha_i) m_{oo} + (-1)^{i+1} \delta\alpha_3 AT(\beta_3) m_{ik(i)} \right) + \int X(\delta) m_{oo} ,$$

$$(5.26) \quad \Theta_1(M_{NN}) = \sum_{i=1}^2 \int \left(-T_i(\delta\beta_0) m_{io} + (-1)^{i+1} AT(\beta_3) m_{ik(i)} \right) + \int \Theta_1(\delta) m_{oo} ,$$

$$(5.27) \quad \Theta_2(M_{NN}) = \sum_{i=1}^2 \int \left(-T_i(\delta\beta_3) m_{oo} + T_i(\delta\beta_3) m_{io} + (-1)^i \delta\beta_3 AT(\beta_3) m_{ik(i)} \right) + \int \Theta_2(\delta) m_{oo} .$$

Proof: The meaning of (3.8) in Corollary 3.1 is that $M_{T_i T_j} = 0$ and hence, (5.21) holds. The Transversality Condition satisfied by m means that

$$(5.28) \quad \nabla_{T_1} m(T_1, T_j) + \nabla_{T_2} m(T_2, T_j) + \nabla_N m(N, T_j) = 0 ,$$

$$(5.29) \quad \nabla_{T_1} m(T_1, N) + \nabla_{T_2} m(T_2, N) + \nabla_N m(N, N) = 0 .$$

We begin by computing (5.22); we have

$$X(M_{T_j N}) = \int \delta X m(T_j, N) + \int X(\delta) m(T_j, N) .$$

Now using integration by parts (taking into account that m is compactly supported in $\bar{\Omega}$), Lemma 5.2, (5.28) and taking $k(j) \neq j \in \{1, 2\}$, we obtain

$$\begin{aligned} \int \delta X m(T_j, N) &= \sum_{i=1}^2 \int \delta\alpha_i T_i m_{jo} + \int \delta\alpha_3 N m_{jo} \\ &= \sum_{i=1}^2 \int -T_i(\delta\alpha_i) m_{jo} + \int \delta\alpha_3 \nabla_N m(T_j, N) + \delta\alpha_3 m(\nabla_N T_j, N) \\ &= \sum_{i=1}^2 \int \left(-T_i(\delta\alpha_i) m_{jo} - \delta\alpha_3 T_i m_{ij} + \delta\alpha_3 m(T_i, \nabla_{T_i} T_j) \right) \\ &\quad + \int \delta\alpha_3 m(\nabla_N T_j, N) \\ &= \sum_{i=1}^2 \int \left(-T_i(\delta\alpha_i) m_{jo} + T_i(\delta\alpha_3) m_{ij} + (-1)^j \delta\alpha_3 AT(\beta_3) m_{io} \right) \\ &\quad + \int (-1)^j \delta\alpha_3 AT(\beta_3) m_{k(j)o} . \end{aligned}$$

To establish (5.23), we note that

$$\Theta_1(M_{T_j N}) = \int \delta \Theta_1 m(T_j, N) + \int \Theta_1(\delta) m(T_j, N) .$$

It follows from similar arguments that

$$\begin{aligned} \int \delta \beta_0 \nabla_N m(T_j, N) + \delta \beta_0 m(\nabla_N T_j, N) &= \\ &= \sum_{i=1}^2 \int -\delta \beta_0 \nabla_{T_i} m(T_i, T_j) + \int \delta \beta_0 m(\nabla_N T_j, N) \\ &= \sum_{i=1}^2 \int \left(-\delta \beta_0 T_i m_{ij} + \delta \beta_0 m(T_i, \nabla_{T_i} T_j) \right) + \int (-1)^j \delta \beta_0 AT(\beta_3) m_{k(j)o} \\ &= \sum_{i=1}^2 \int \left(T_i(\delta \beta_0) m_{ij} + (-1)^j \delta \beta_0 AT(\beta_3) m_{io} \right) + \int (-1)^j \delta \beta_0 AT(\beta_3) m_{k(j)o} . \end{aligned}$$

The remaining formulas follow from similar computations. ■

Proof of Theorem 1.1 with Condition II: It follows from Proposition 5.1 that there is a constant $C_1 > 0$ such that

$$(5.30) \quad \|dM\|_0 \leq C_1 \varepsilon \|m\|_0 ,$$

where ε can be made arbitrarily small by requiring that g be sufficiently close to the euclidean metric. On the other hand, by the Poincaré inequality, we obtain that there is a constant $C_2 > 0$ such that

$$(5.31) \quad \|H\|_1 \leq C_2 \|dH\|_0 ,$$

and using Lemma 5.1 with $n = 3$, it follows that there is a constant $C_3 > 0$ such that

$$(5.32) \quad \|m\|_0 \leq C_3 \|M\|_1 .$$

Using (5.30)–(5.32) we get that $M = 0$ and, consequently, $m = 0$. ■

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